

Tidal Disruption Events in the TDAMM era

Lessons learned from ZTF

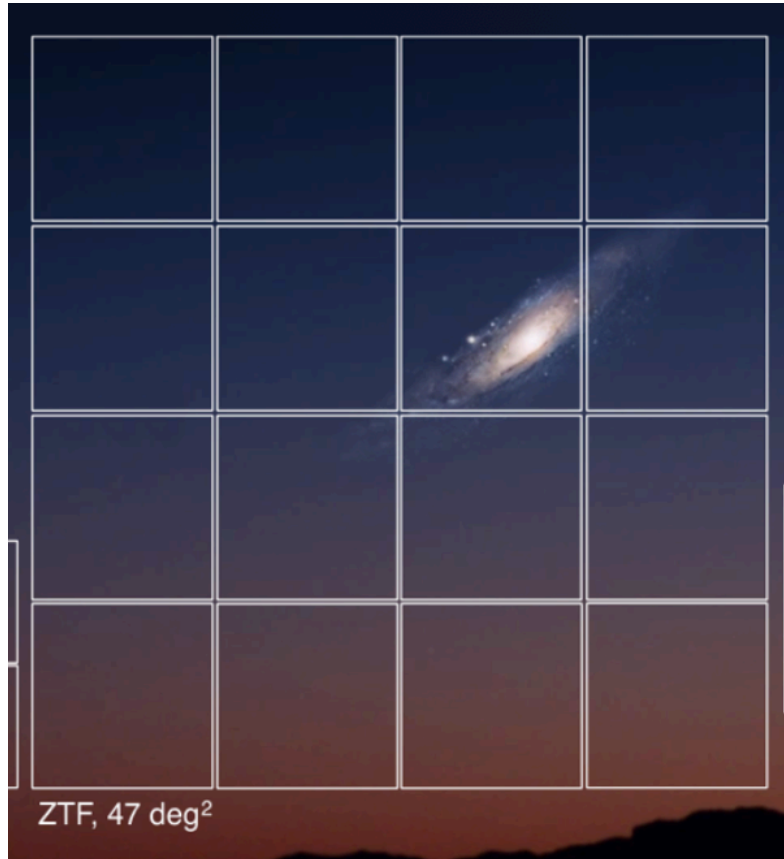
Robert Stein

Postdoctoral Scholar at Caltech

TDAMM Workshop, Annapolis, US

How can we probe the multi-messenger emission of TDEs?

Introducing the Zwicky Transient Facility



Credit: Iair Arcavi

ZTF is an optical telescope with a 47 sq. deg. field of view

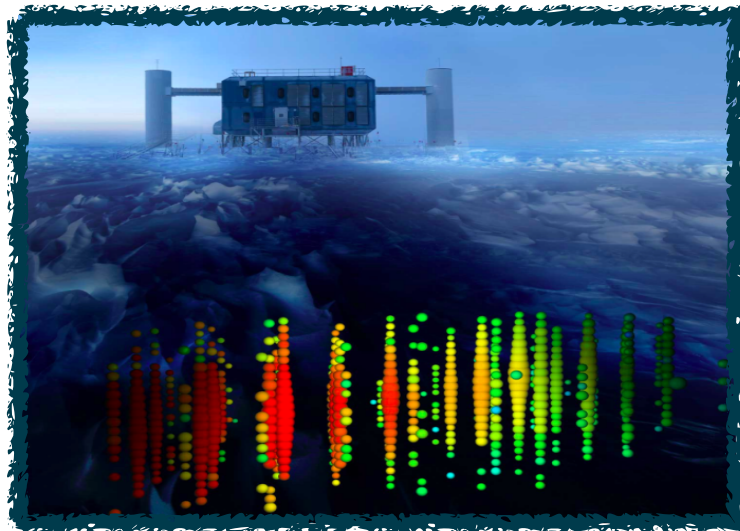
Surveys the entire northern sky every 2 nights, in g+r, as part of a public survey

The ZTF neutrino follow-up program

Stein et al. 2022



Credit: Iair Arcavi



4 Credit: IceCube

~500k objects
per night



Neutrino
direction and
time

Find
counterpart?

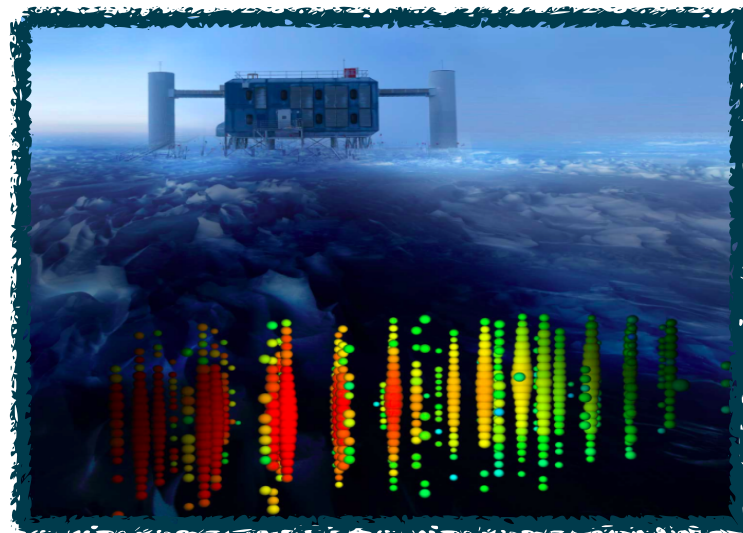
Trigger
Follow-Up

The ZTF neutrino follow-up program

Stein et al. 2022



Credit: Iair Arcavi



5 Credit: IceCube

~500k objects
per night

Neutrino
direction and
time



Reject stars,
asteroids,
planets

Correlate

Spectroscopically
classify few
remaining objects

Find
counterpart?

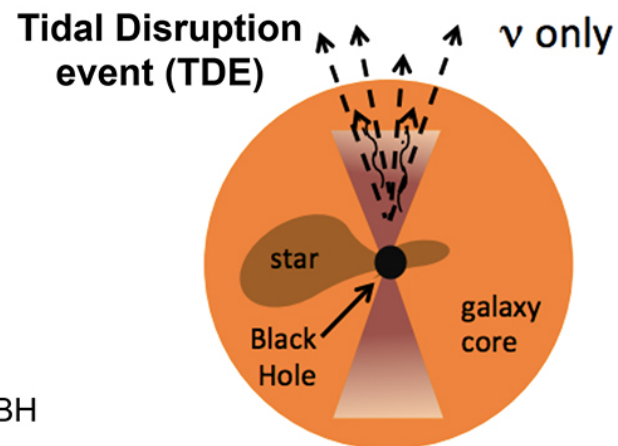
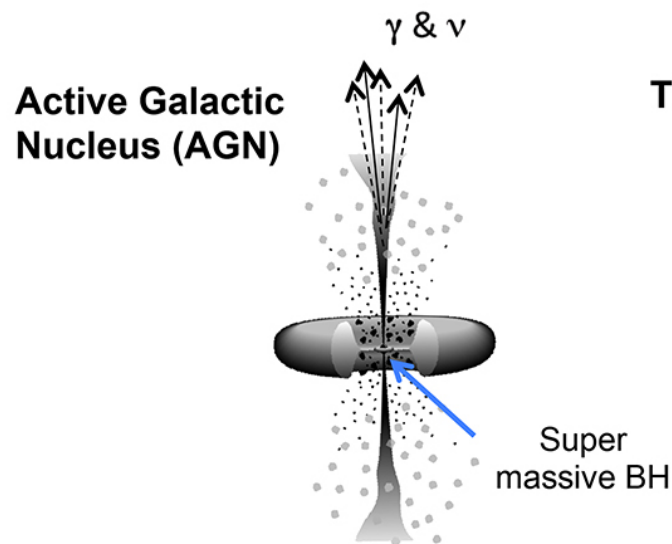
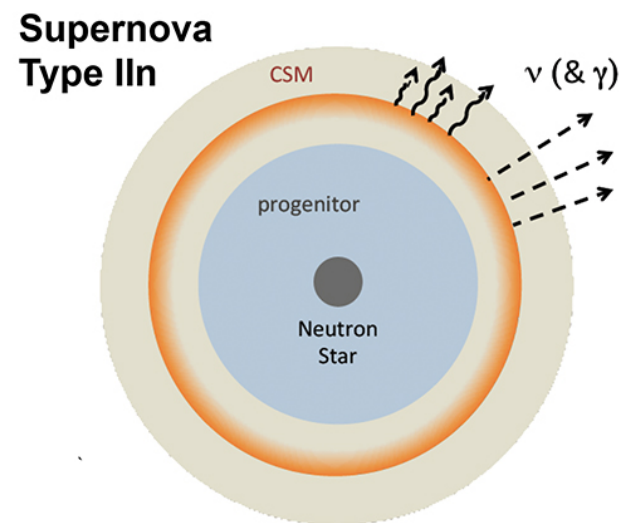
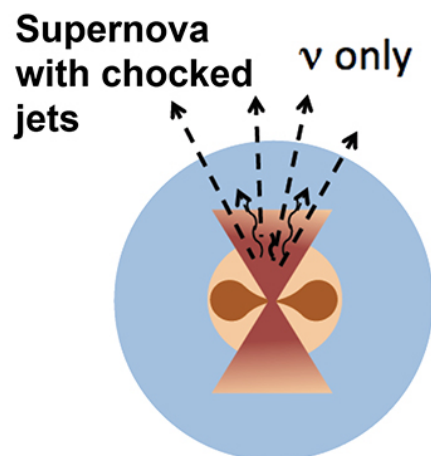
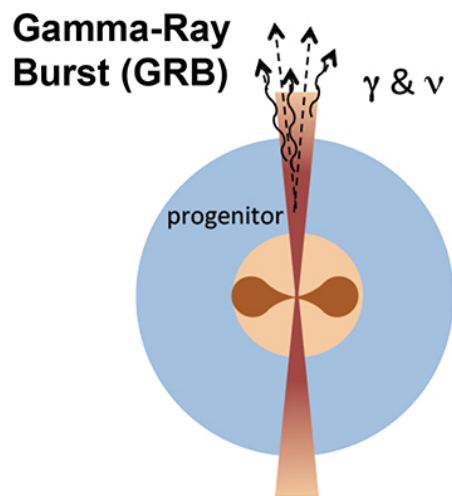
Trigger
Follow-Up



Credit: Laurie Hatch

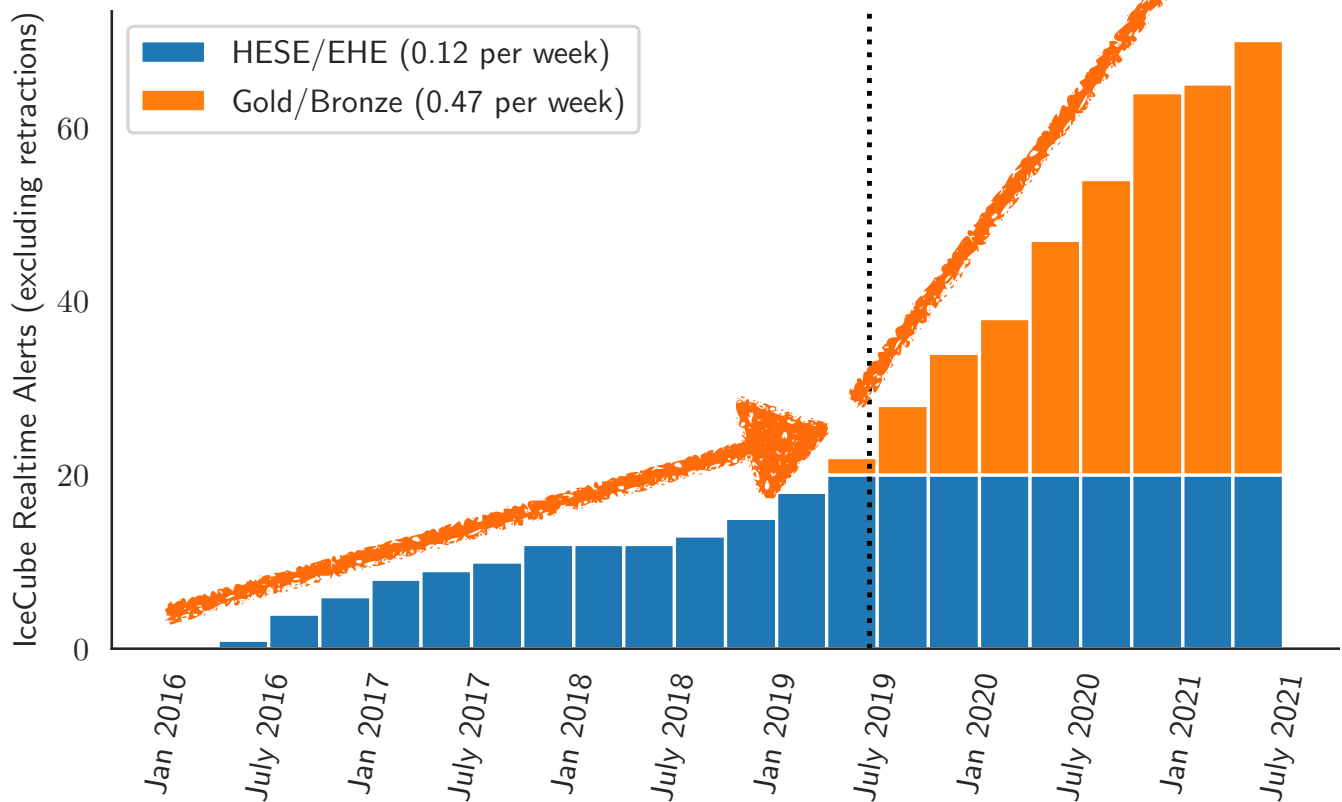
What are we looking for?

Bartos and Kowalski 2017

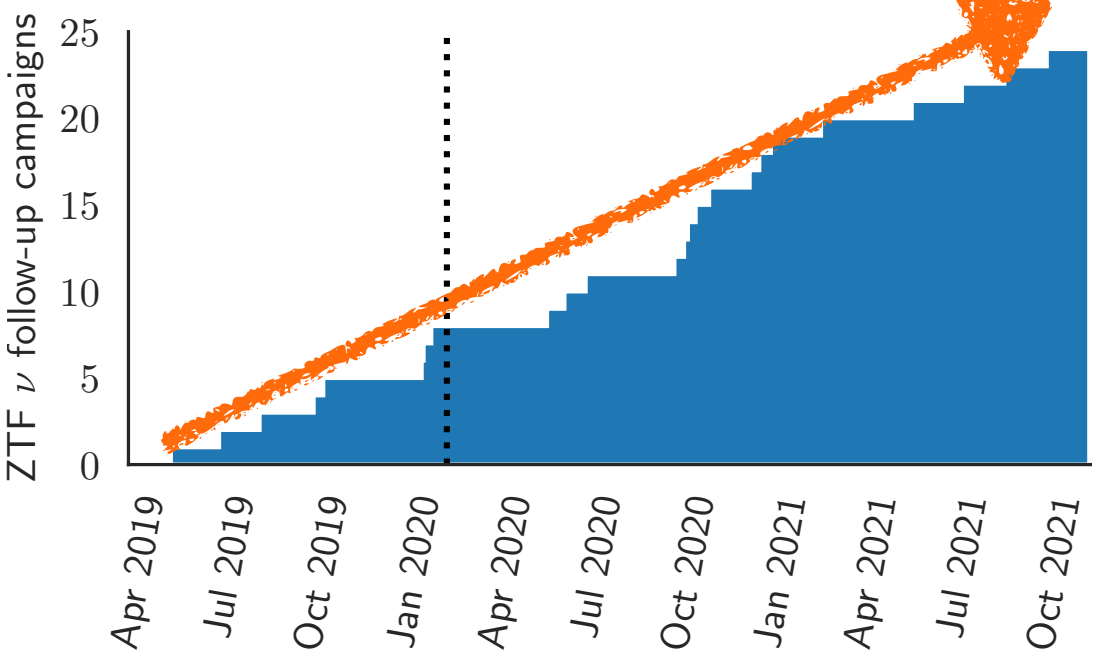


Some statistics...

1 neutrino alert per 2 weeks



1 ZTF campaign per 5 weeks



Stein et al. 2022

nature astronomy

New sources
of neutrinos



Two TDEs coincident with neutrinos

ARTICLES

<https://doi.org/10.1038/s41550-020-01295-8>

nature
astronomy



A tidal disruption event coincident with a high-energy neutrino

Stein et al. 2021

Featured in Physics

Editors' Suggestion

Candidate Tidal Disruption Event AT2019fdr Coincident with a High-Energy Neutrino

Simeon Reusch *et al.*

Phys. Rev. Lett. **128**, 221101 – Published 3 June 2022

Physics See Focus story: [Neutrinos from a Black Hole Snack](#)

What did we learn from ZTF?

Neutrino follow-up works as a method -> Using ~dozens of multi-wavelength follow-up campaigns, you can probe a ~10%-level contribution to the astrophysical neutrino flux.

Performing realtime searches enables prompt and flexible multi-wavelength observations to characterise individual sources.

Theoretical models, informed by both population data and this single-object data, are compatible with neutrino emission from TDEs.

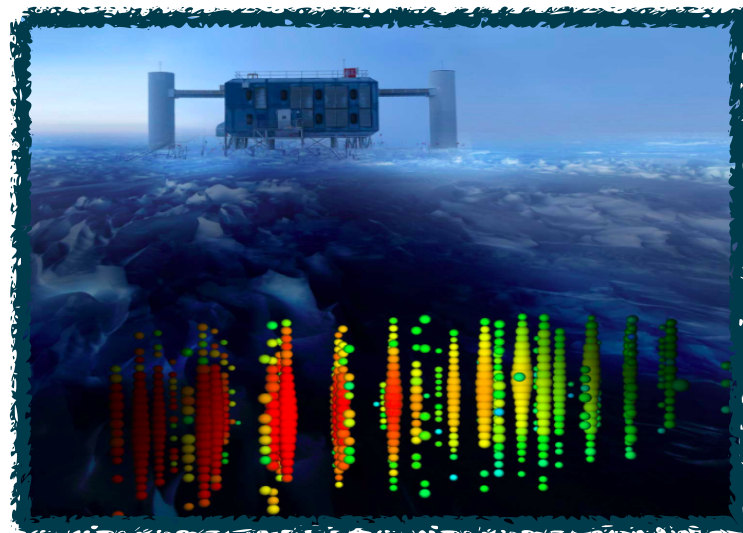
**What do we need to enable
multi-messenger follow-up?**

The ZTF neutrino follow-up program

Stein et al. 2022



Credit: Iair Arcavi



~500k objects
per night

Neutrino
direction and
time



Reject stars,
asteroids,
planets

Correlate



Credit: Laurie Hatch

Spectroscopically
classify few
remaining objects

Find
counterpart?

Trigger
Follow-Up

Any neutrino follow-up program

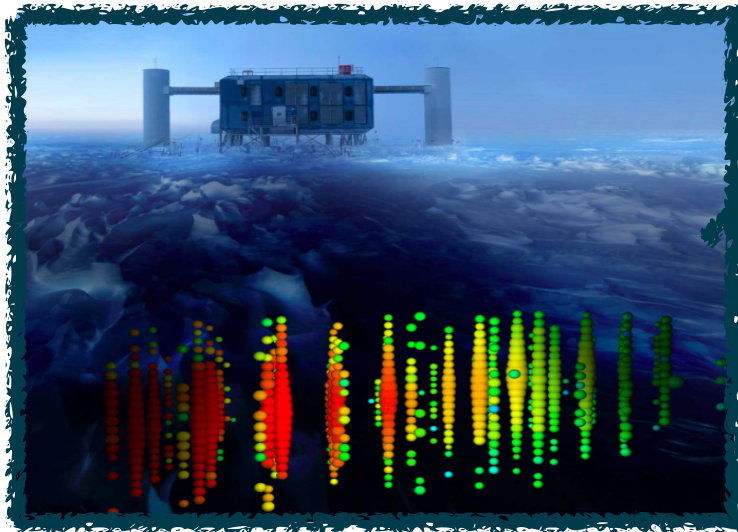
Stein et al. 2022



Telescope to
find
candidates

Pipeline to select candidates that are:

- Spatially/temporally coincident
- Likely astrophysical
- Belong to populations which could emit neutrinos



Neutrino
direction and
time



Spectroscopically
classify few
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Any neutrino follow-up program

Stein et al. 2022

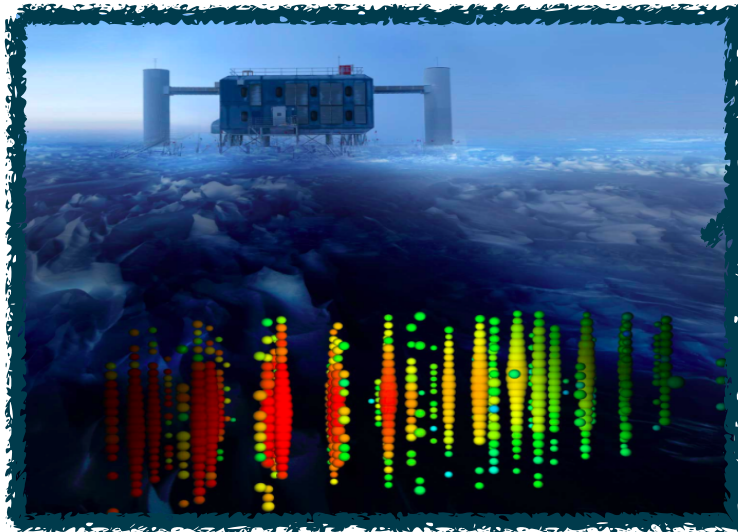


Telescope to
find

Pipeline to select candidates that are:

- Spatially/temporally coincident

1: Prompt and reliable neutrino information is the prerequisite for all downstream multi-messenger analysis. Need accurate localisations, and useful information to prioritise “better” (=more astrophysical) neutrino alerts.



Neutrino
direction and
time



Spectroscopically
classify few
remaining objects

Find
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Follow-Up

Any neutrino follow-up program

Stein et al. 2022

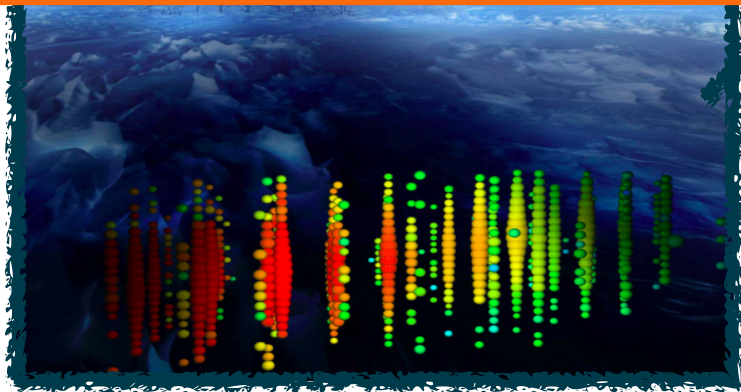


Telescope to
find
candidates

Pipeline to select candidates that are:

- Spatially/temporally coincident
- Likely astrophysical
- Belong to populations which could emit neutrinos

2. Once accounting for detector systematics, neutrino “90% contours” will almost always be a few square degrees. Suitably wide-field telescopes are required to tile these regions and identify possible EM counterparts.



Neutrino
direction and
time

Find
counterpart?

Trigger
Follow-Up

Any neutrino follow-up program

Stein et al. 2022

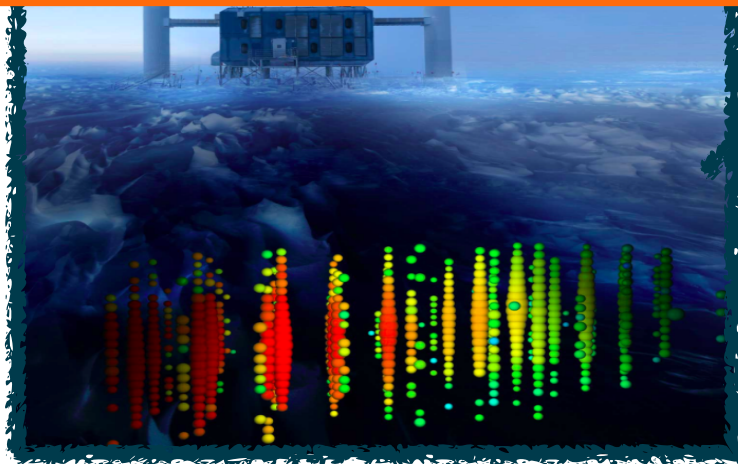


Telescope to
find
candidates

Pipeline to select candidates that are:

- Spatially/temporally coincident
- Likely astrophysical
- Belong to populations which could emit neutrinos

3. Pipelines are needed to select candidates. This is primarily done through algorithmic filtering, ML, catalogue matching. Public survey data is invaluable here.



Neutrino
direction and
time

Credit: Laurie Hatch

Find
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Any neutrino follow-up program

Stein et al. 2022



Telescope to
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Pipeline to select candidates that are:

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4. Need dedicated spectroscopic programs to support EM follow-up.

time



Spectroscopically
classify few
remaining objects

Trigger
Follow-Up

Any neutrino follow-up program

Stein et al. 2022

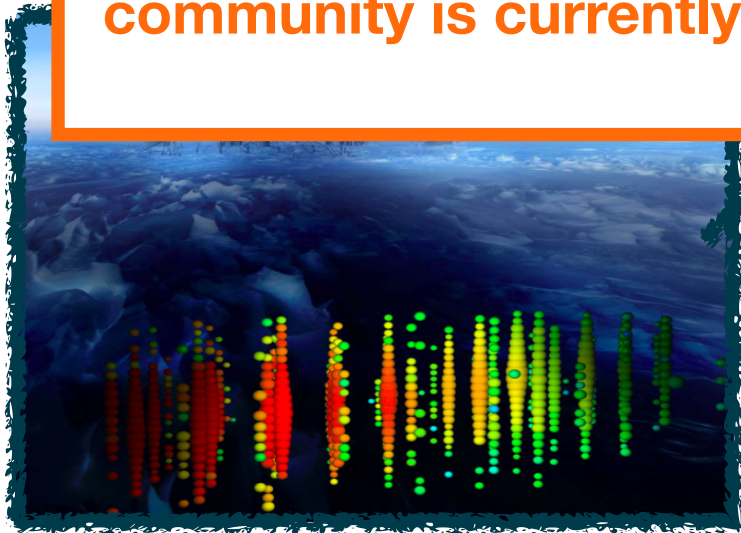


Telescope to
find
candidates

Pipeline to select candidates that are:

- Spatially/temporally coincident
- Likely astrophysical
- Belong to populations which could emit neutrinos

5. We require multi-wavelength observations to characterise the sources. These could be from dedicated proposals, but DDTs have also worked well. The community is currently operating a “kitchen sink” model, and that will presumably still be true in another decade.



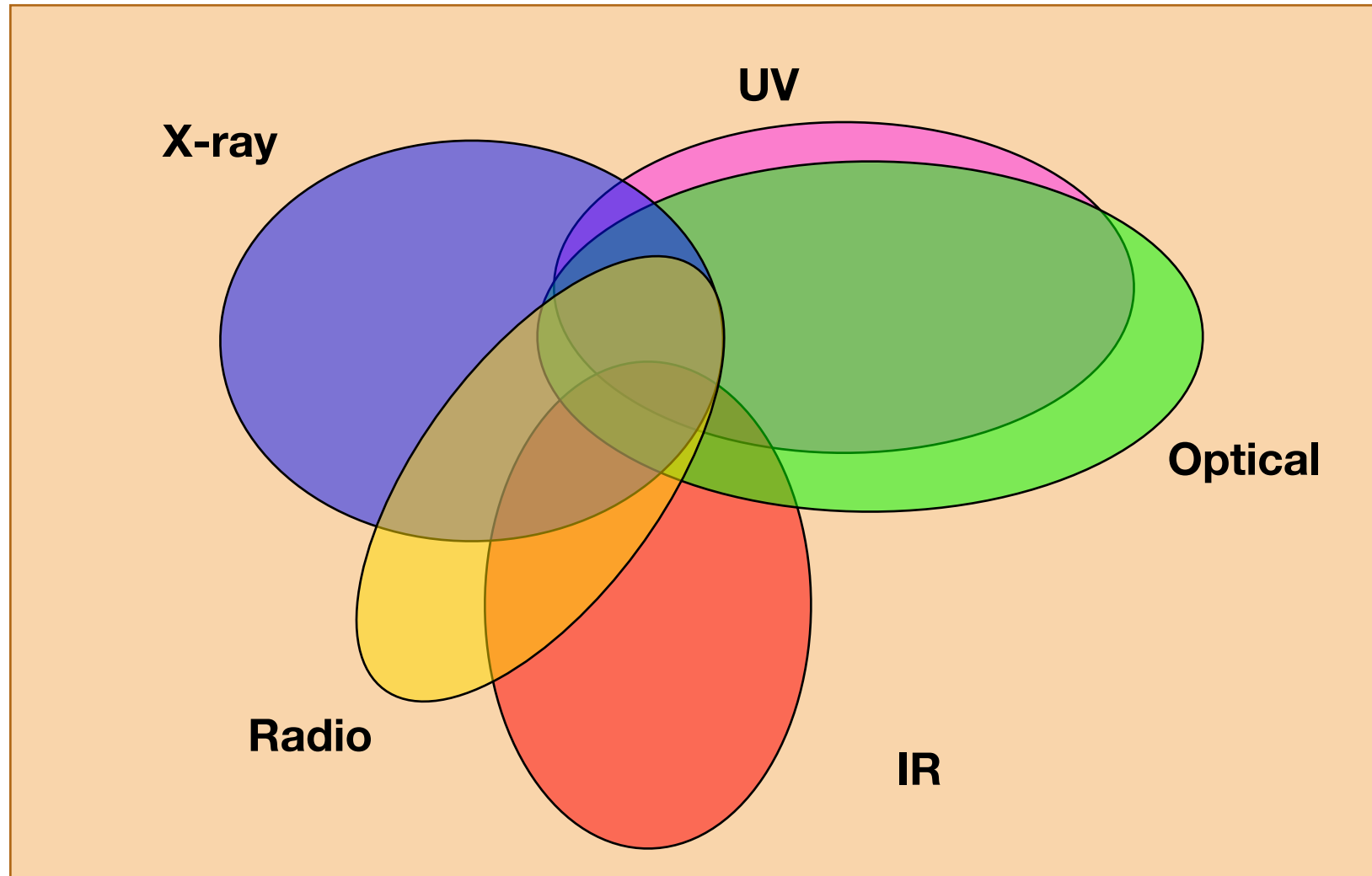
Neutrino
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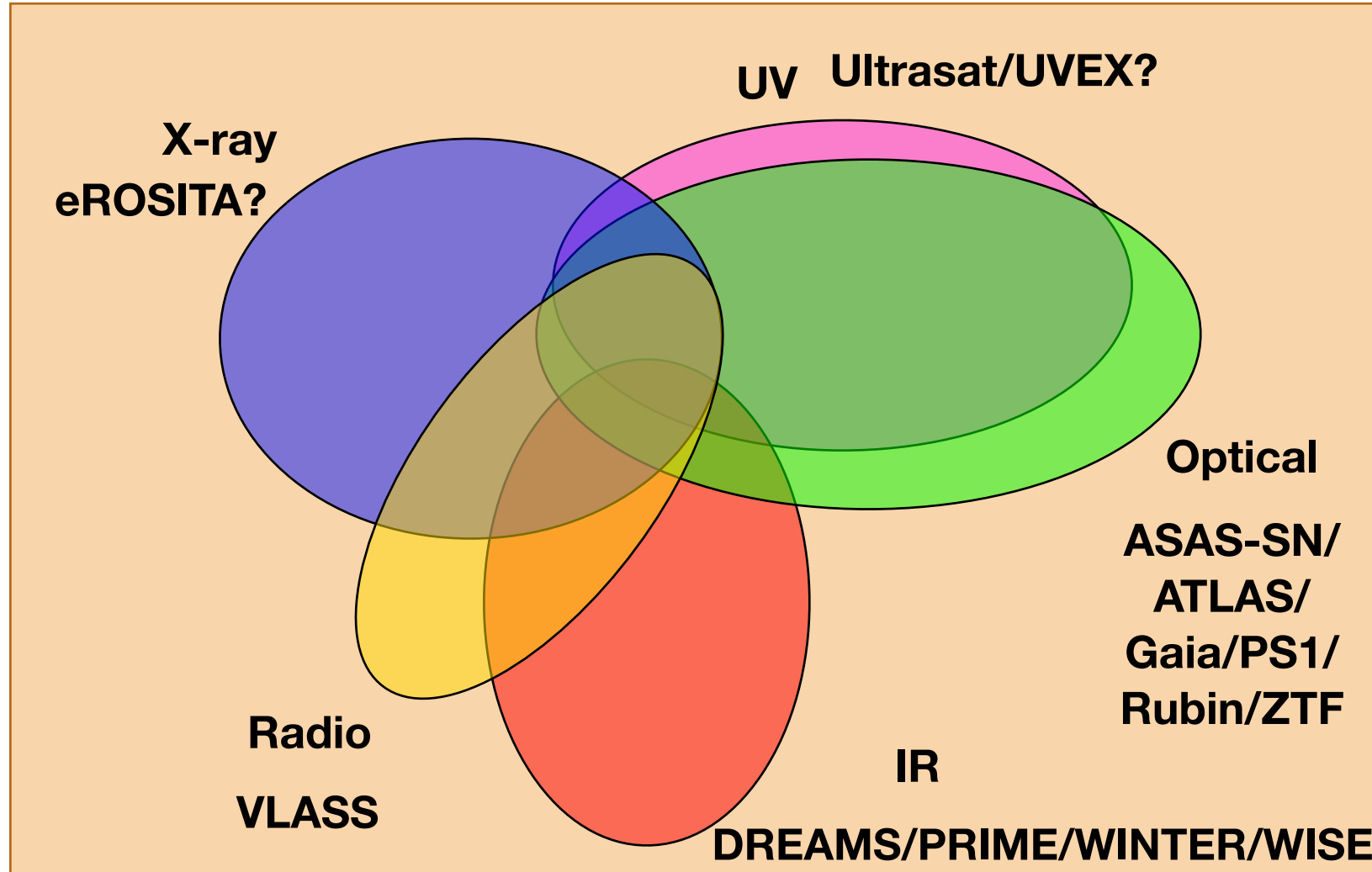
Beyond Optical Follow-Up

Towards a pan-chromatic view of TDEs



Not to scale (yet!)

Towards a pan-chromatic view of TDEs



Not to scale (yet!)

**Systematic searches
are possible. Need
wide-field surveys for
each wavelength!**

WINTER: Wide-field InfraRed Transient Explorer



Credit: Robert Stein

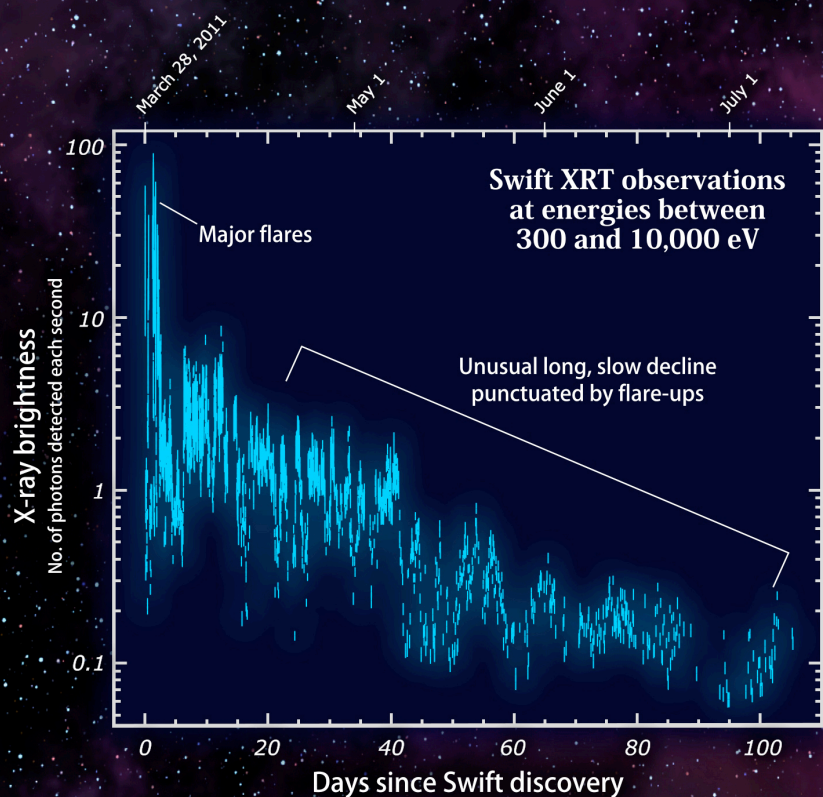
WINTER will do a J-band survey at ~monthly cadence, AND dedicated neutrino follow-up!

Thinking outside the box

The view from 2011

The New York Times

X-rays from Swift J1644+57



Credit: NASA

OBSERVATORY

In Celestial Twist, Black Hole Swallows a Dying Star

By Sindya N. Bhanoo

June 16, 2011

In what sounds like a one of a kind murder mystery, a dying star has fallen into a black hole and been ripped apart.

The event, which was observed on March 28, was originally thought to be a gamma ray burst from a collapsing star, but researchers suspected something more sinister was at play. Their findings appear in a pair of papers published online by the journal Science.

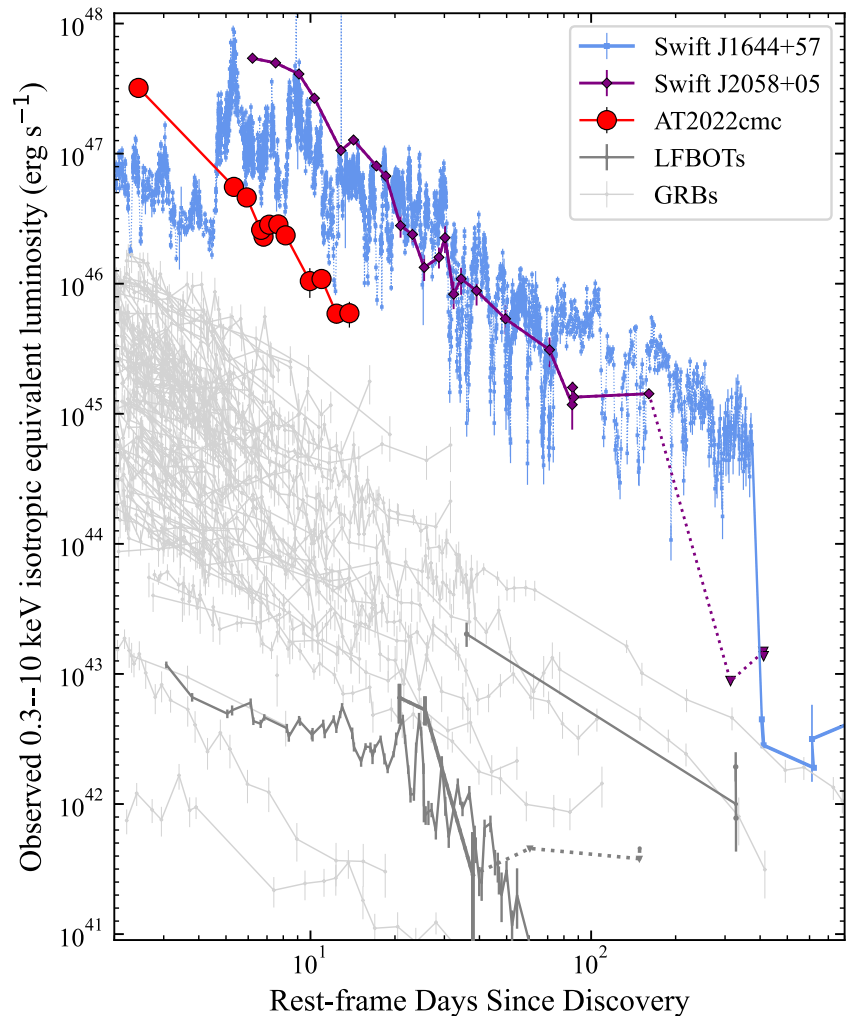
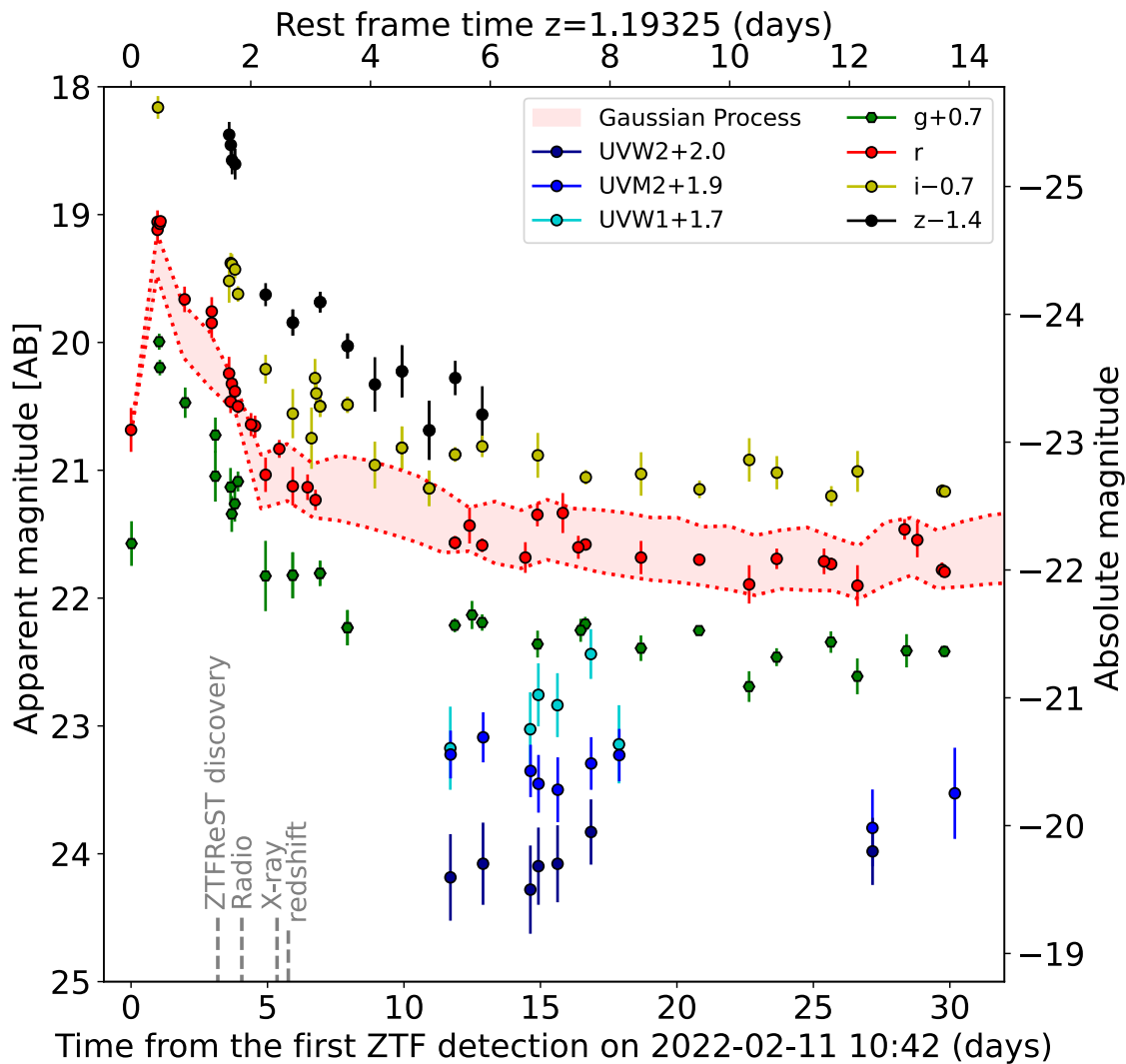
Jetted TDEs are “the dog that didn’t bark”

TDEs in the TDAMM era | Robert Stein | TDAMM | 8/23/22

Caltech

Back to the future: An optically-selected on-axis “jetted TDE”

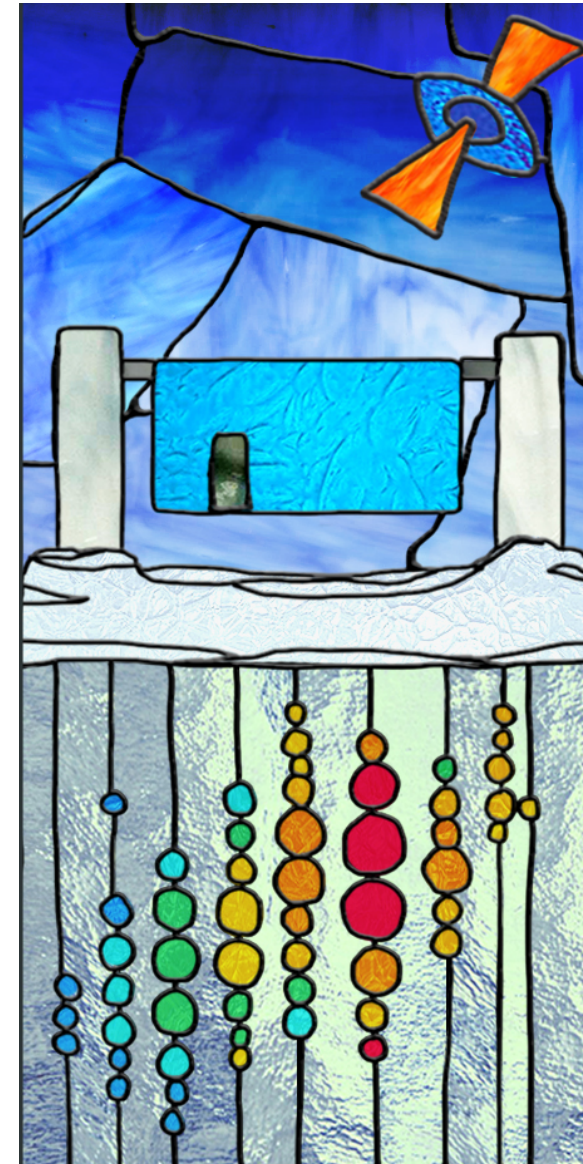
Andreoni et al.
submitted



Summary

Summary

- TDE science has had a great decade, let's hope the next one is just as fruitful!
- Neutrino follow-up can probe the multi-messenger emission of TDEs, but requires an integrated TDA ecosystem.
- We need to better understand the intersecting multi-wavelength properties of TDEs. Requires wide-field surveys.
- It was an unexpectedly slow decade for relativistic TDEs. Optically-selected AT2022cmc has provided an exciting leap forward, and suggests we need to pay more attention to “rapid transients”.

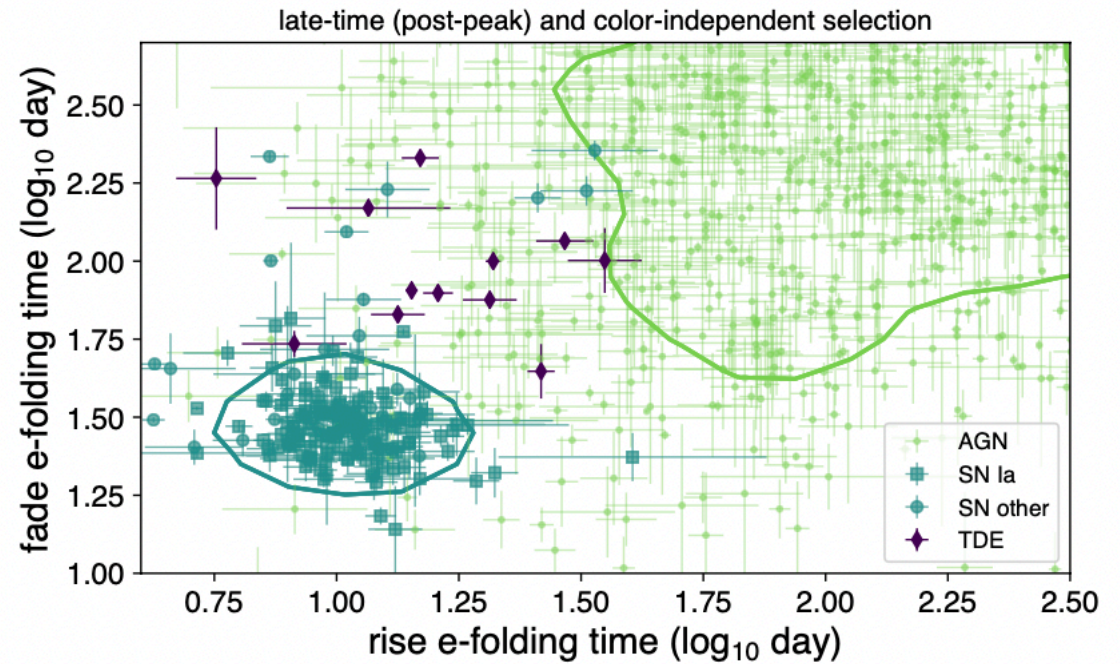
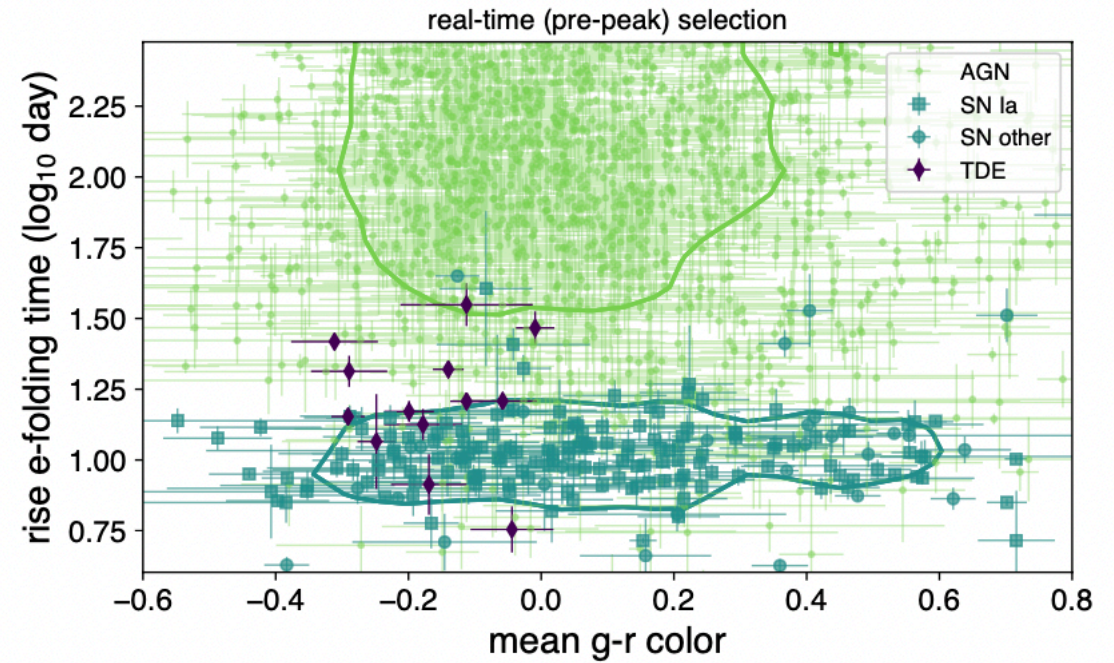
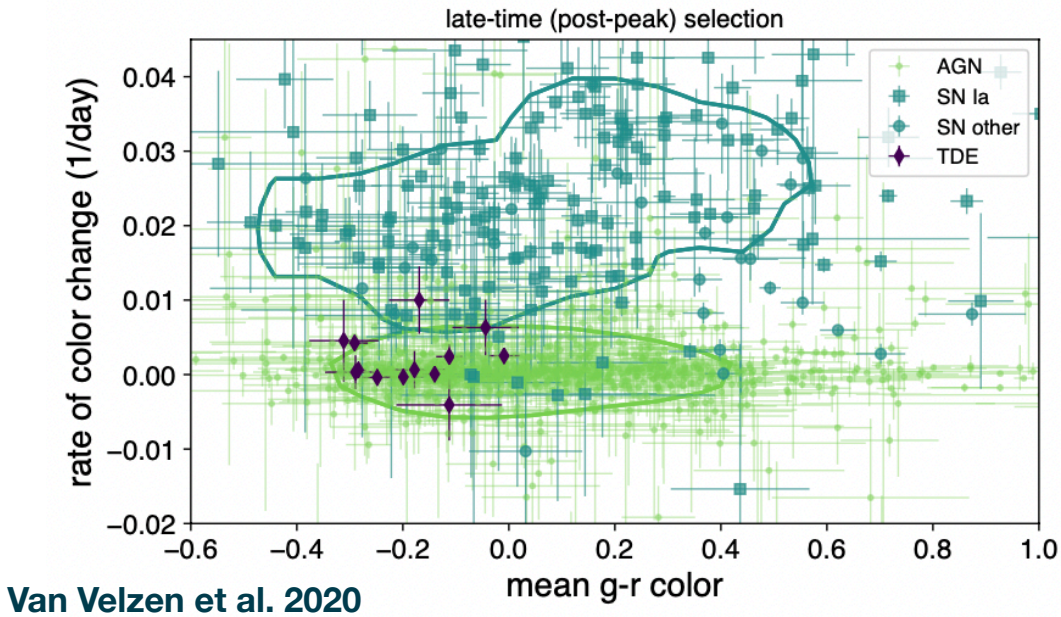


Credit: IceCube

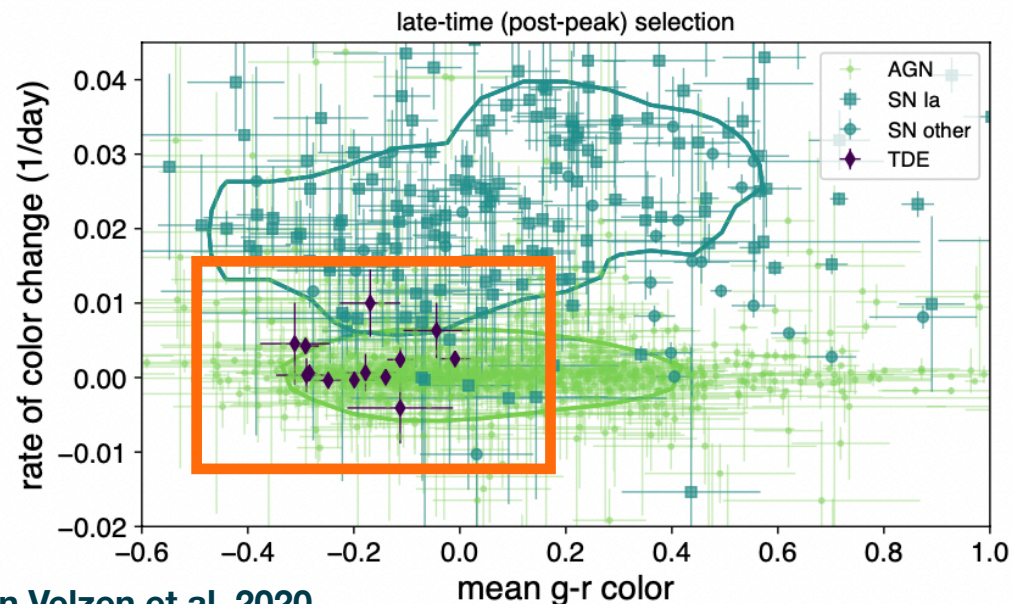
Backup

Embracing diversity in TDEs

How does ZTF find a TDE?

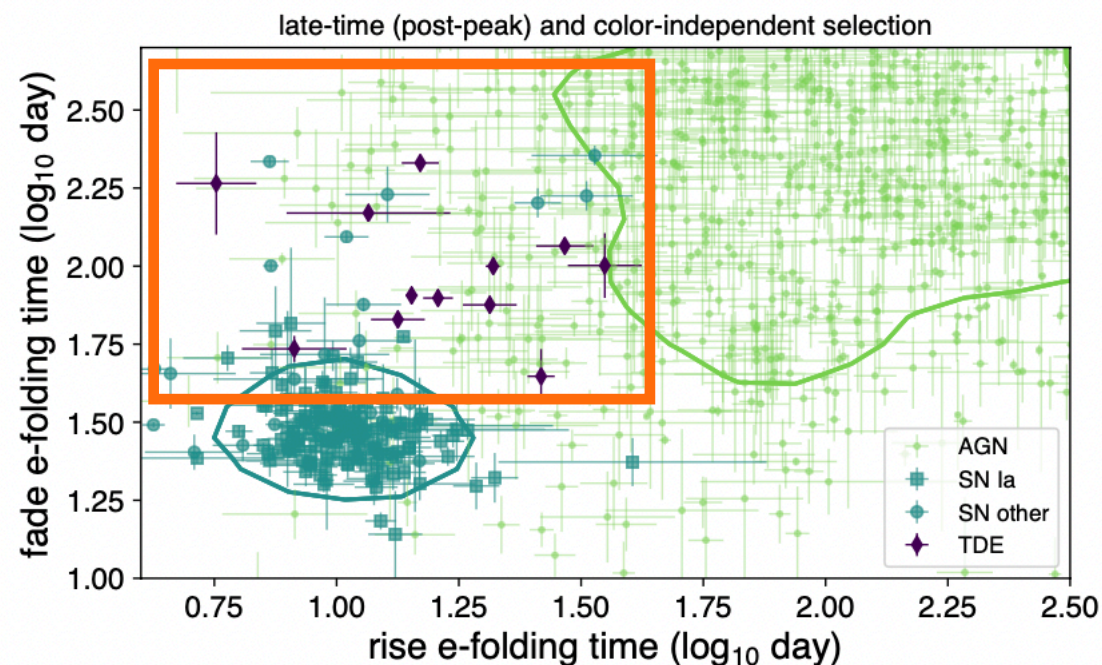
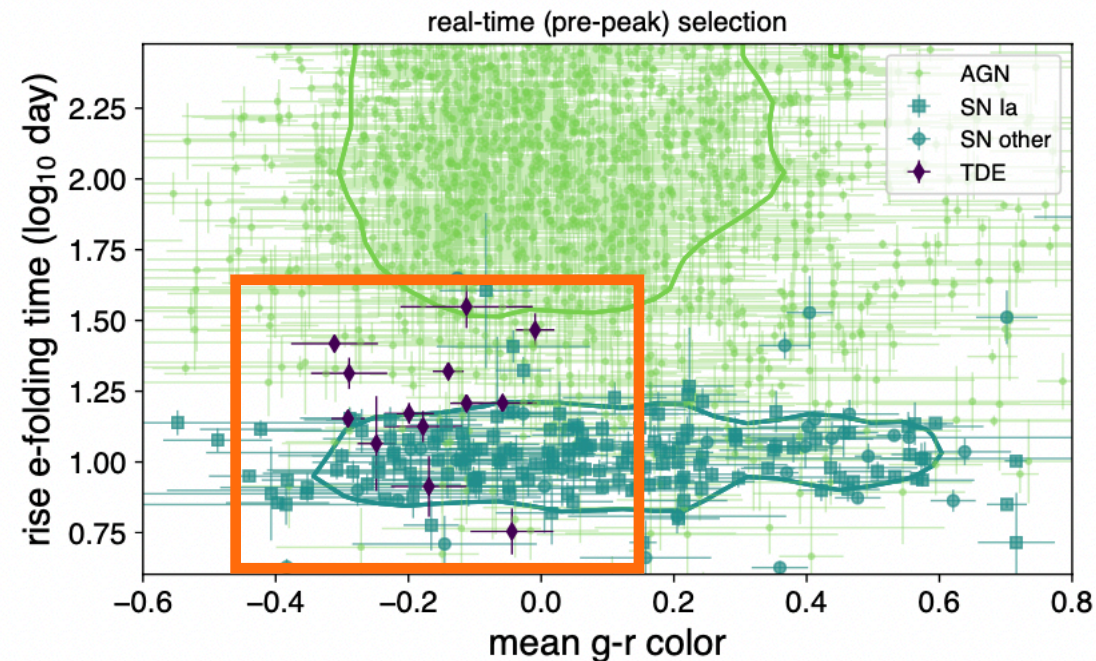


How does ZTF find a TDE?



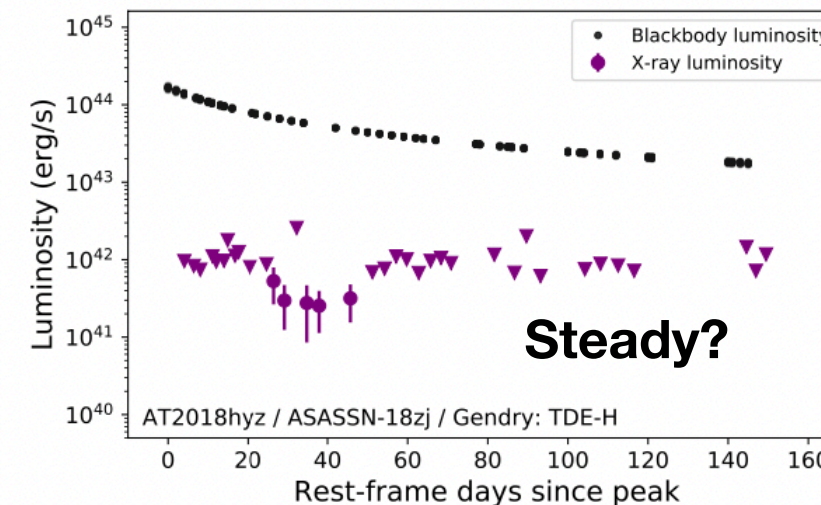
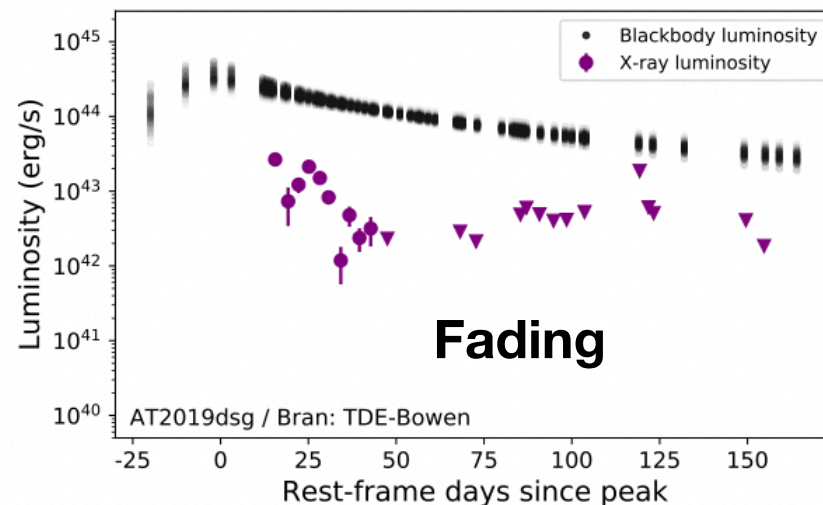
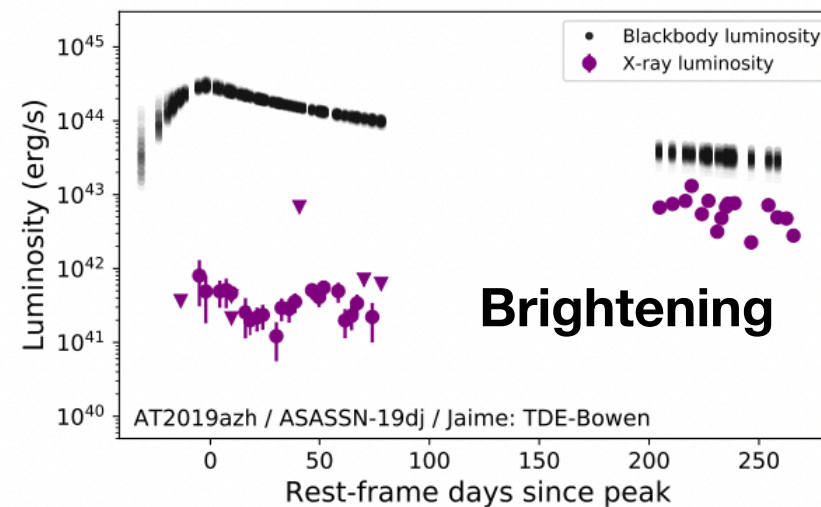
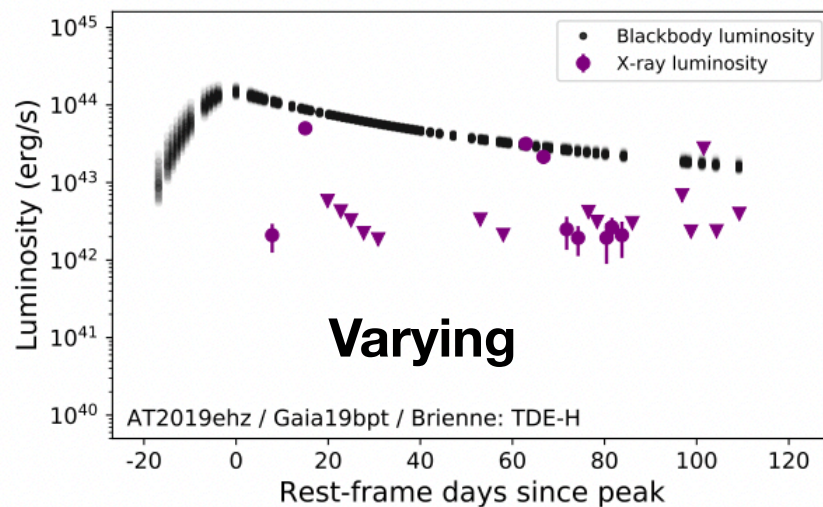
Van Velzen et al. 2020

ZTF operates a “Rubin-style” selection function to select TDE candidates with relatively high purity, yielding a systematic optical sample

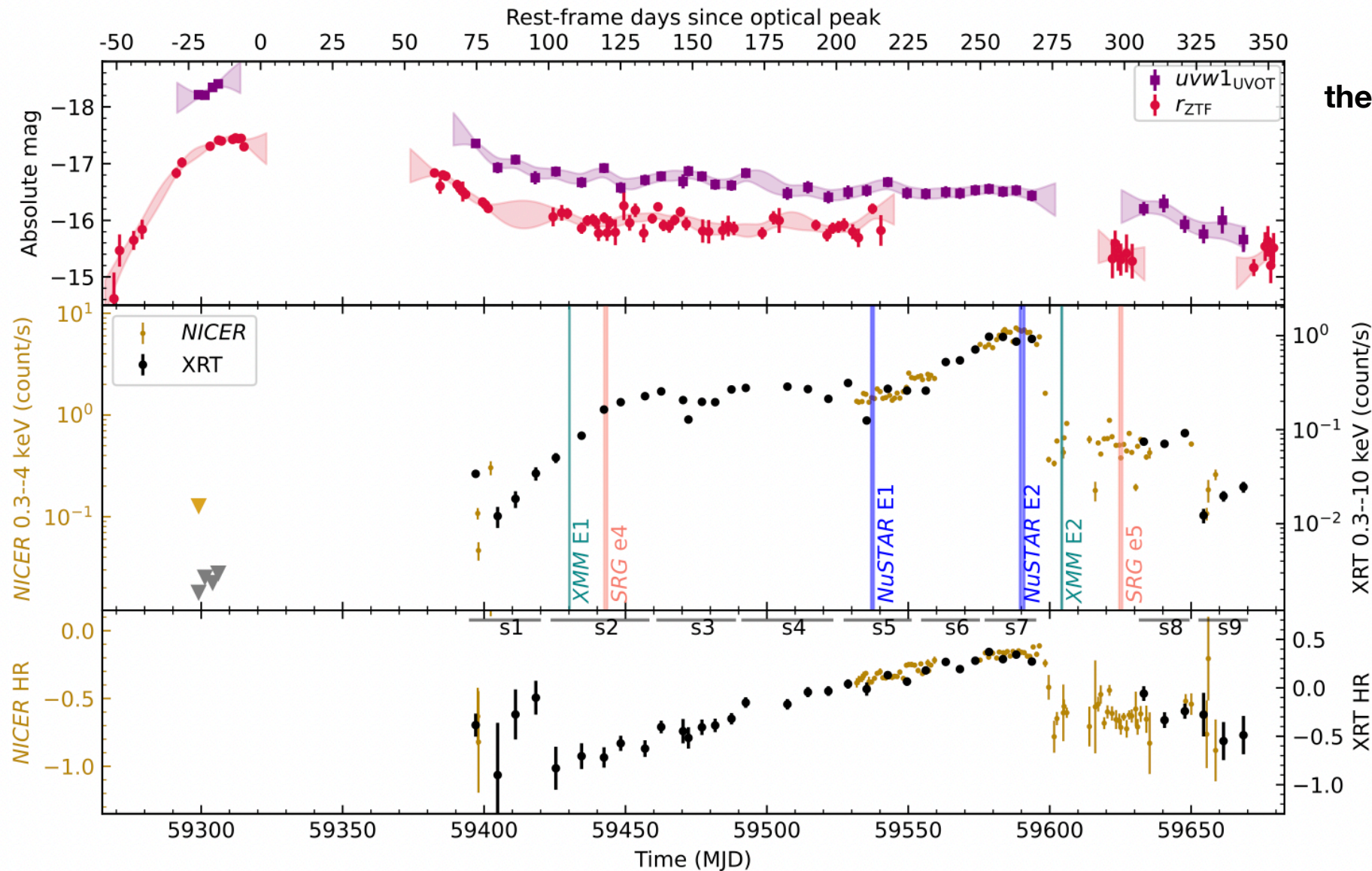


Case Study: X-ray properties of 17 optical TDEs

Van Velzen et al. 2020



More recently: AT2021ehb with dramatic X-ray/HR evolution

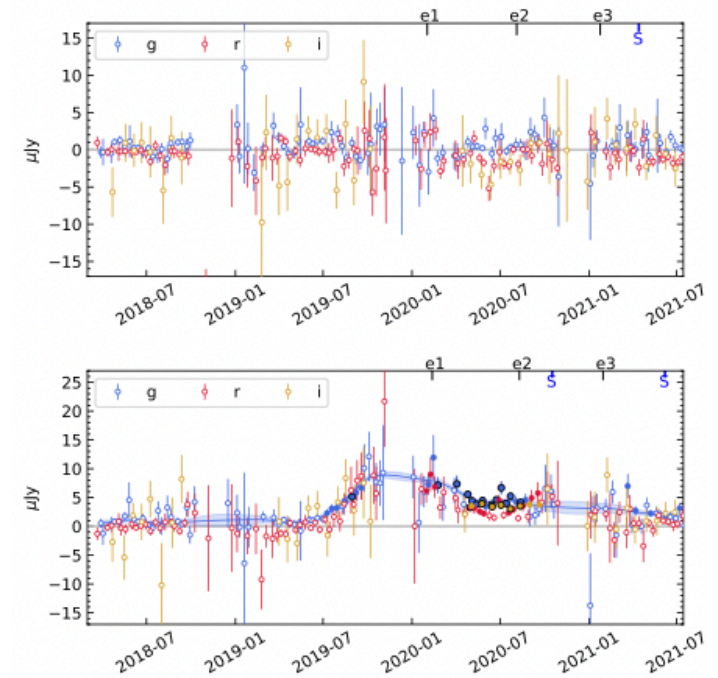
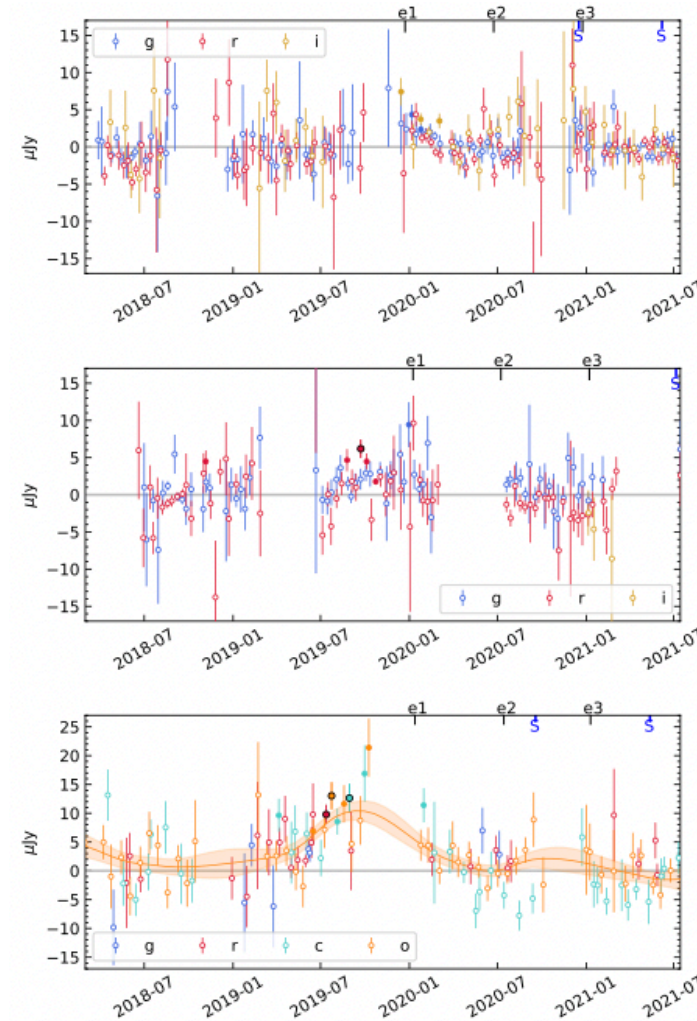


Not just a soft
thermal component!

X-ray-selected TDEs from eROSITA tell a different story

Sazonov et al 2022.

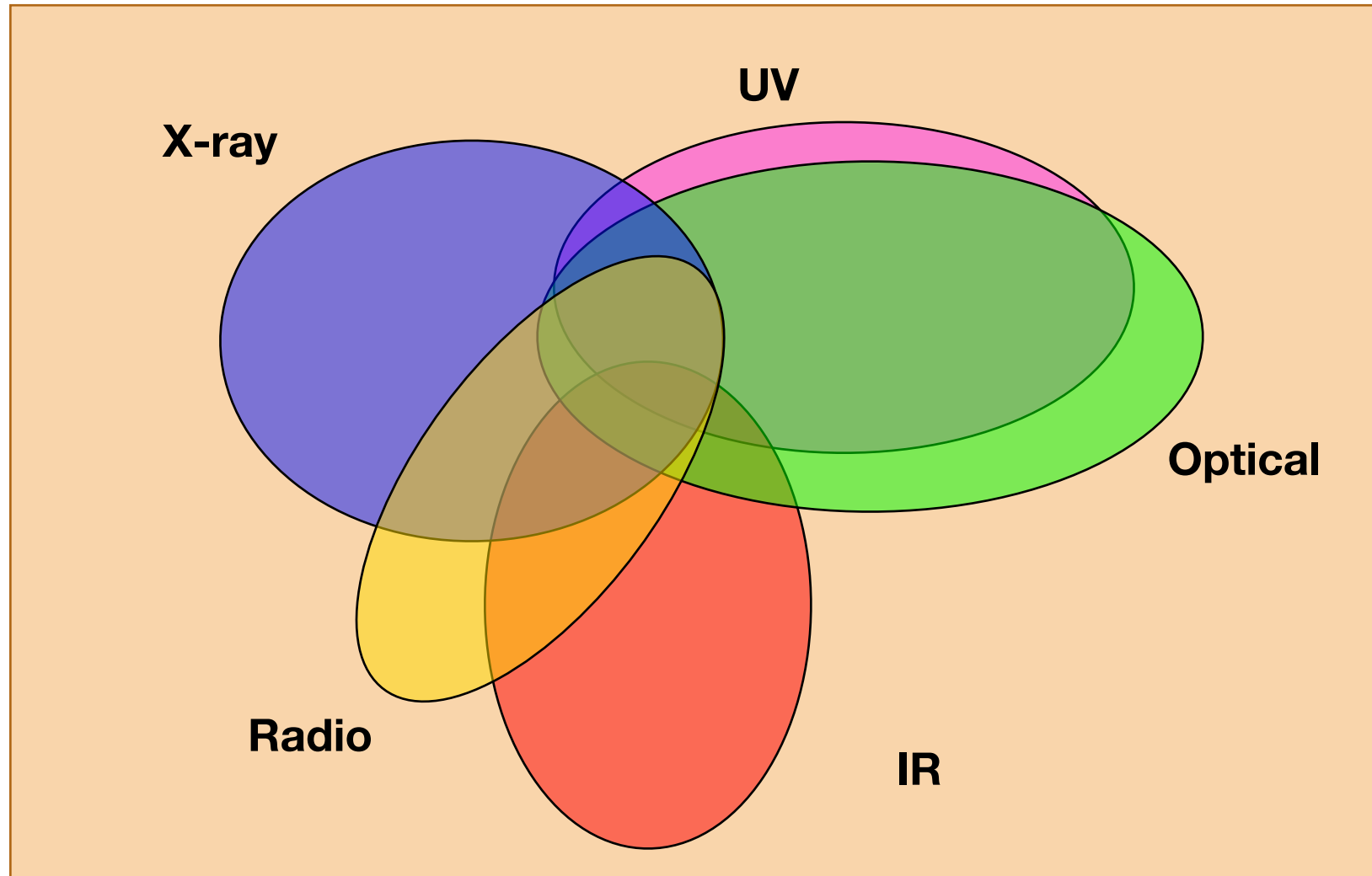
Object (SRGE)	Optical flare peak AB mag
J135514.8+311605	$g > 21.83$
J013204.6+122236	$g > 21.51$
J153503.4+455056	$21.33 < o < 20.58$
J163831.7+534020	$g > 21.84, r > 21.86$
J163030.2+470125	$21.47 < g < 20.71$
J021939.9+361819	$g > 21.43$
J161001.2+330121	$g > 21.68$
J171423.6+085236	$c > 20.04$
J071310.6+725627	$g > 21.41$
J095928.6+643023	$22.40 < r < 21.64$
J091747.6+524821	$21.49 < r < 20.74$
J133053.3+734824	$g > 21.73$
J144738.4+671821	$c > 20.41$



All eROSITA TDEs are optically faint, with $L_{\text{opt}}/L_x < 0.3$

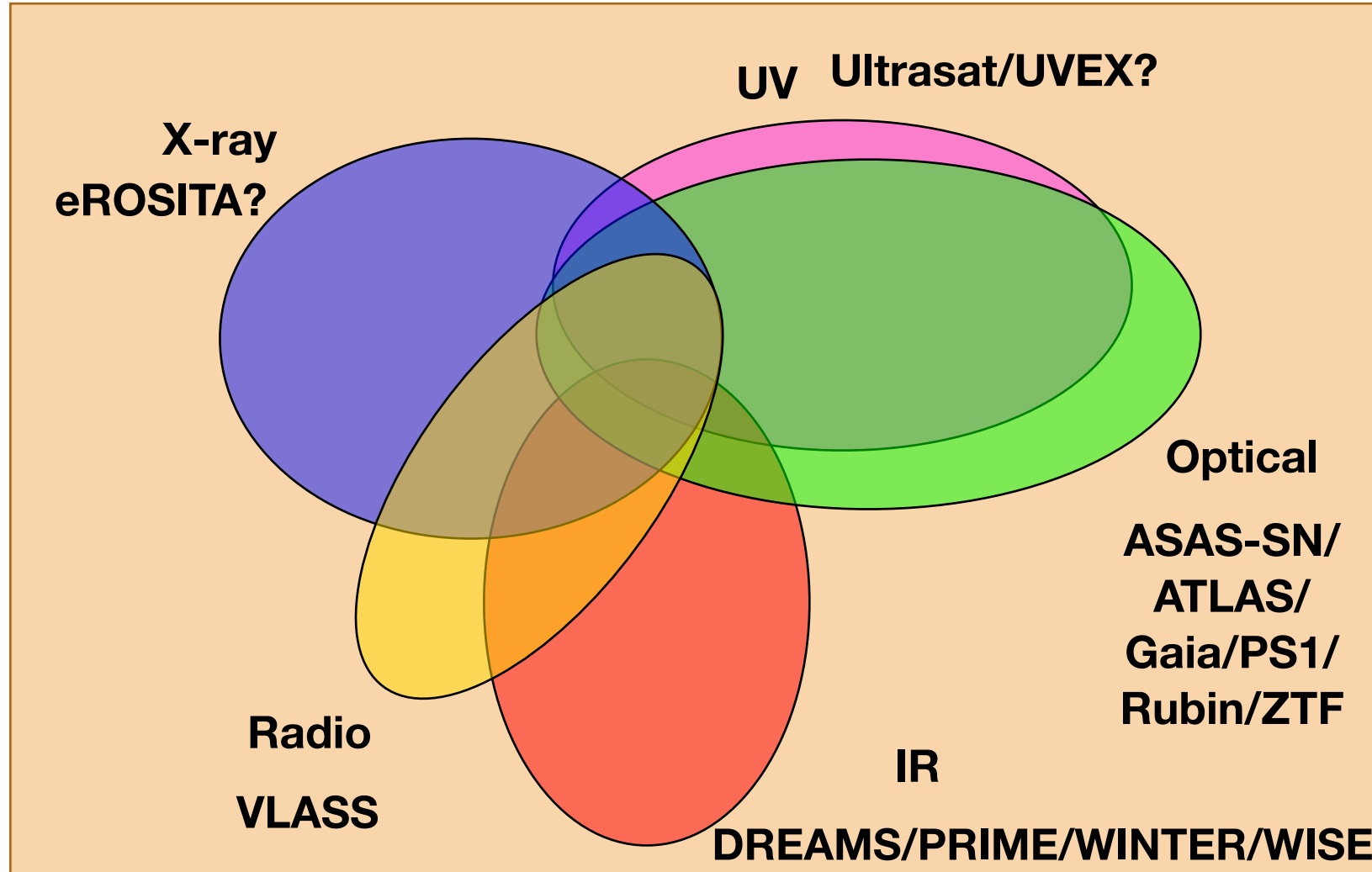
Only 4/13 detected

Towards a pan-chromatic view of TDEs



Not to scale (yet!)

Towards a pan-chromatic view of TDEs



Not to scale (yet!)

Systematic searches are possible. Need wide-field surveys for each wavelength!

WINTER: Wide-field InfraRed Transient Explorer



Credit: Robert Stein

WINTER will do a J-band survey at ~monthly cadence, AND dedicated neutrino follow-up!

Bridging theory and observation

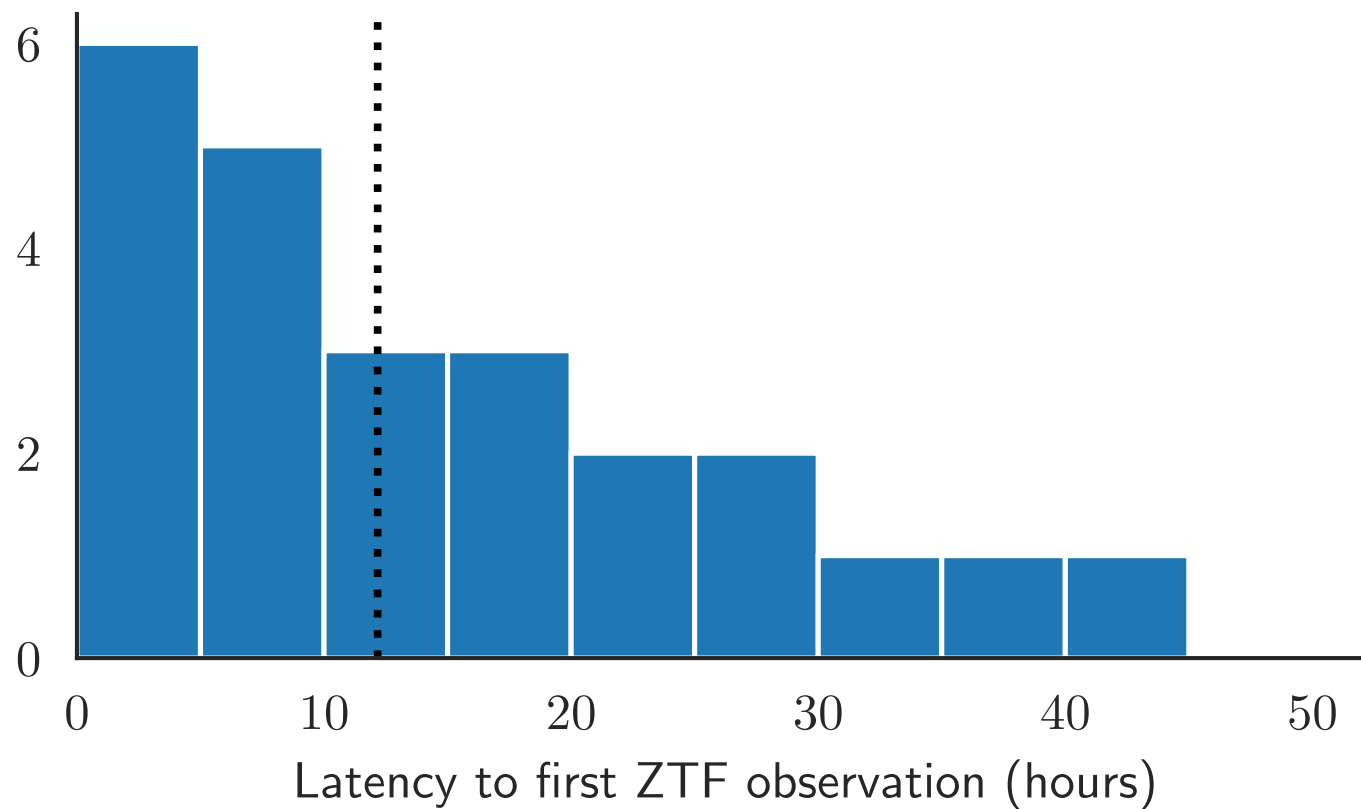
We currently lack information to inform theoretical models of TDE neutrino/CR emission.

Off the top of my head, critical open questions for theorists include:

- **What is the local rate of (all) TDEs, and of jetted TDEs specifically?**
- **How do these rates change as a function of redshift?**
- **What is the distribution of outflows, winds or jets in TDEs? How common are they? How do they vary by object?**
- **Do White-Dwarf/IMBH TDEs exist? How common are they? These in particular are studied for potential UHECR emission, but there is a substantial gap in our experimental knowledge. (We've found ~ 0 thus far.)**
- **How well can we distinguish a TDE in an AGN from an extreme AGN flare? What are the differences (if any) between TDE accretion and AGN accretion?**

Some statistics...

Median latency of 12 hours from neutrino detection to ZTF observation

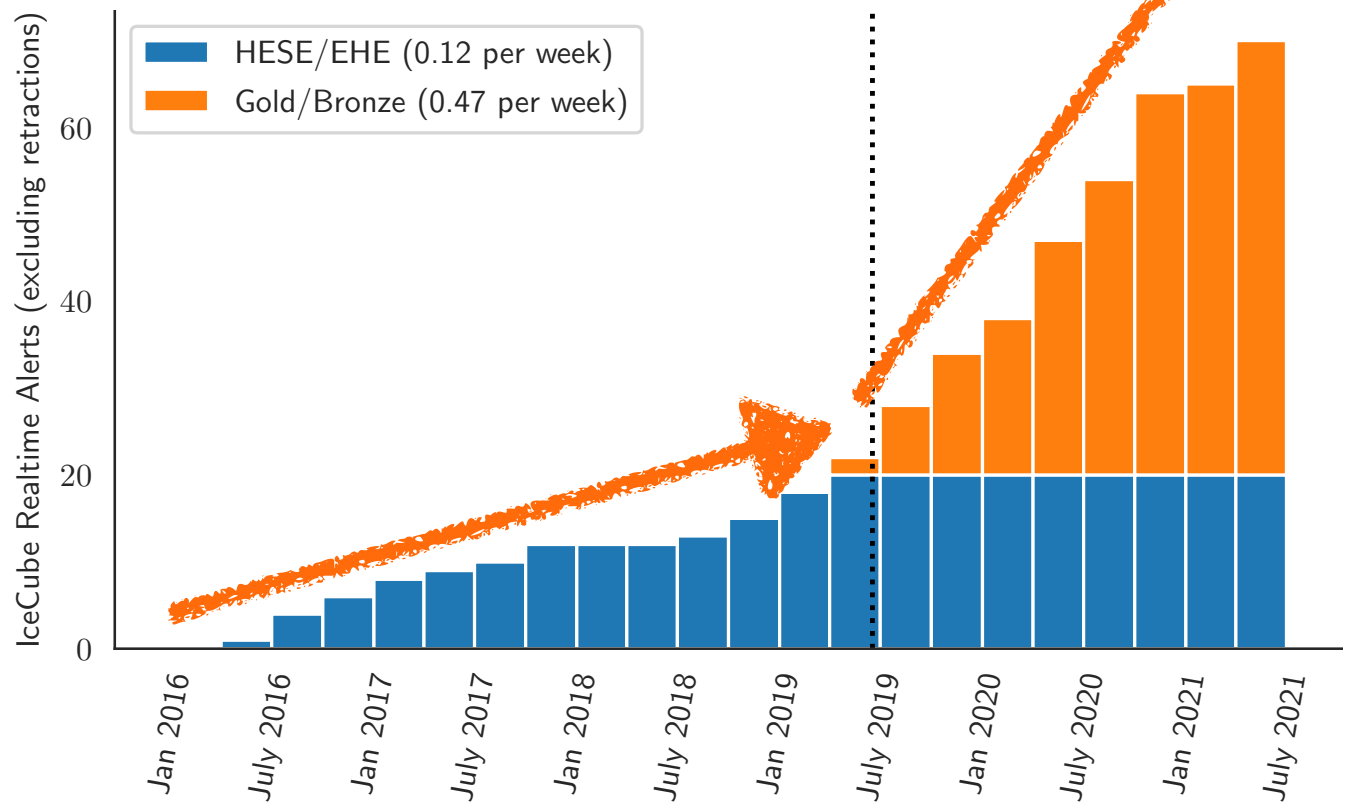


Stein et al. 2022

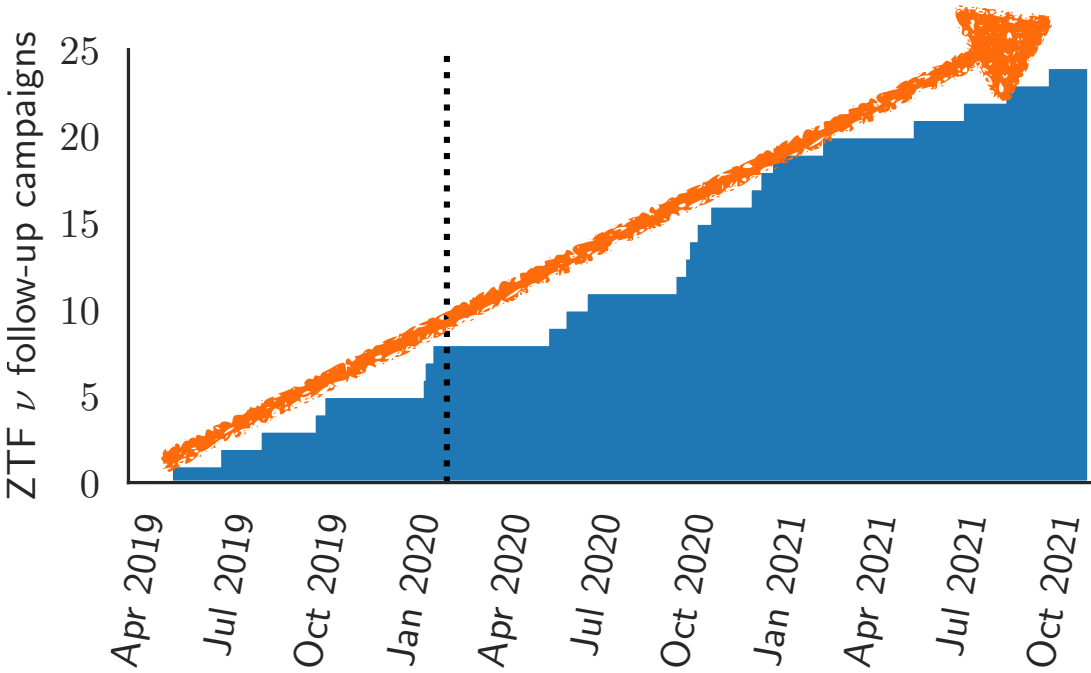
What did we find?

Some statistics...

1 neutrino alert per 2 weeks



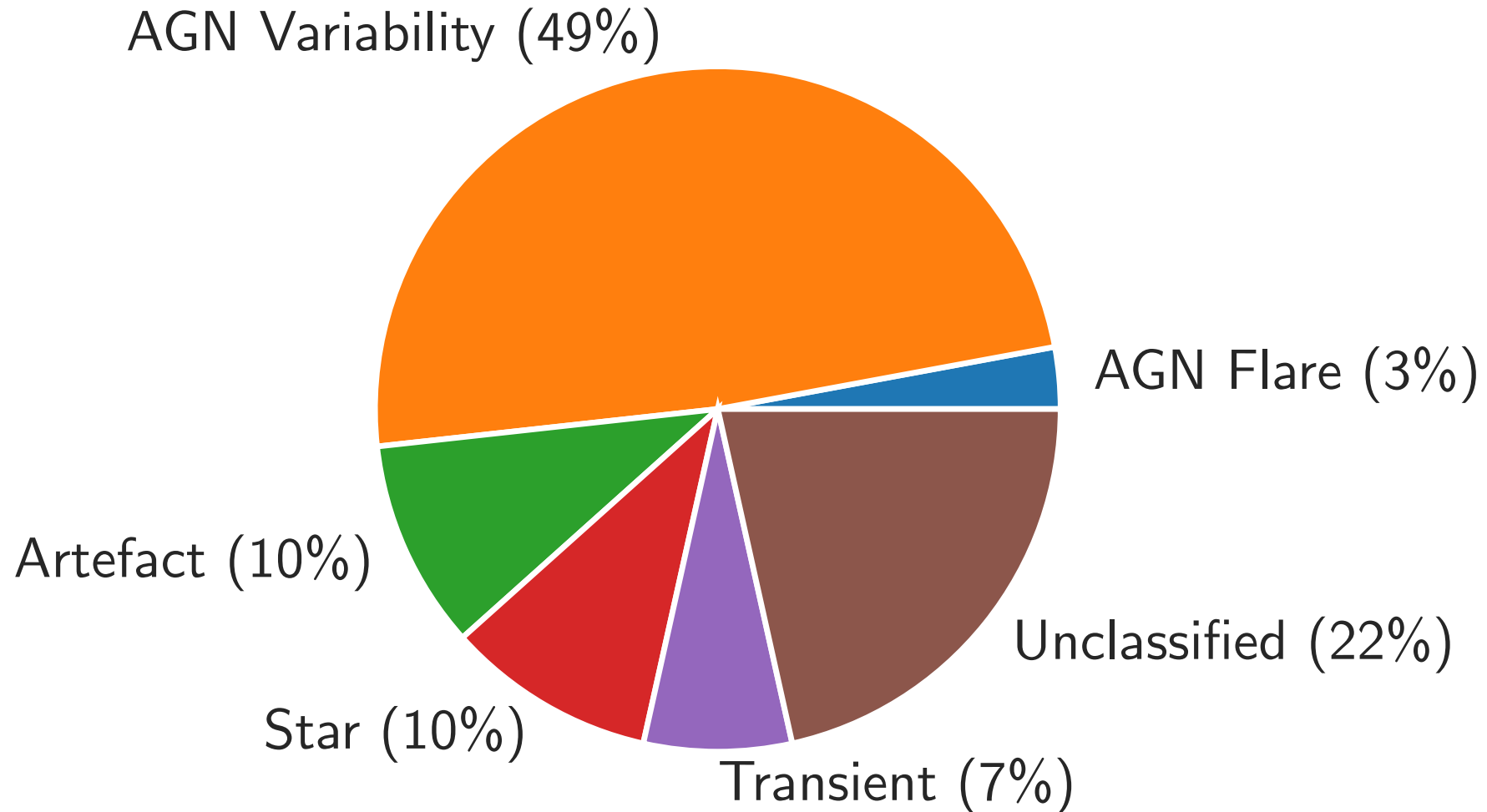
1 ZTF campaign per 5 weeks



Stein et al. 2022

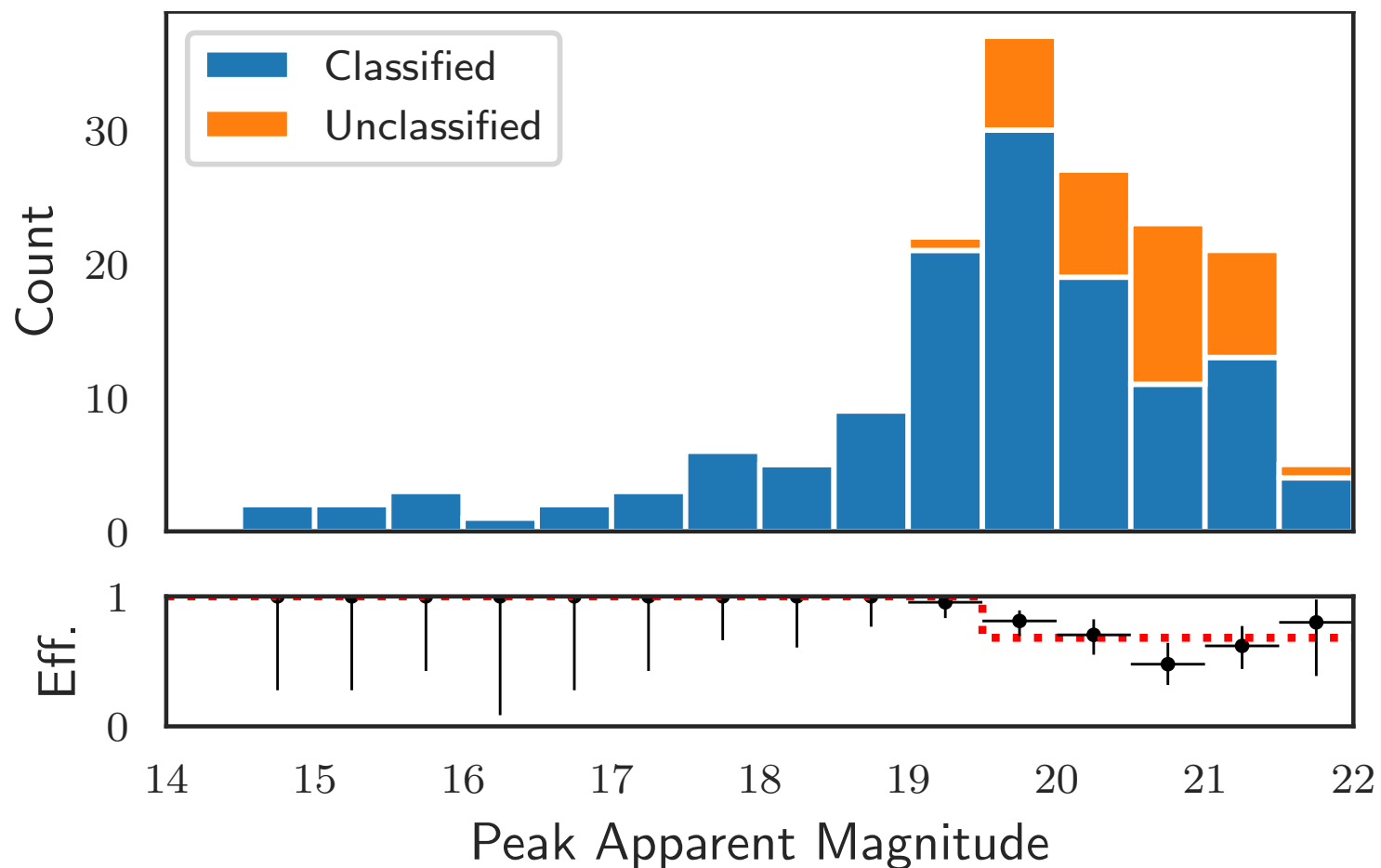
Some statistics...

Only a small fraction of 172 candidates are ultimately transients



Some statistics...

We classify ~100% of bright sources and ~70% of fainter ones



Stein et al. 2022

nature astronomy

New sources
of neutrinos



Two TDEs coincident with neutrinos

ARTICLES

<https://doi.org/10.1038/s41550-020-01295-8>

nature
astronomy

 Check for updates

A tidal disruption event coincident with a high-energy neutrino

Stein et al. 2021

Featured in Physics

Editors' Suggestion

Candidate Tidal Disruption Event AT2019fdr Coincident with a High-Energy Neutrino

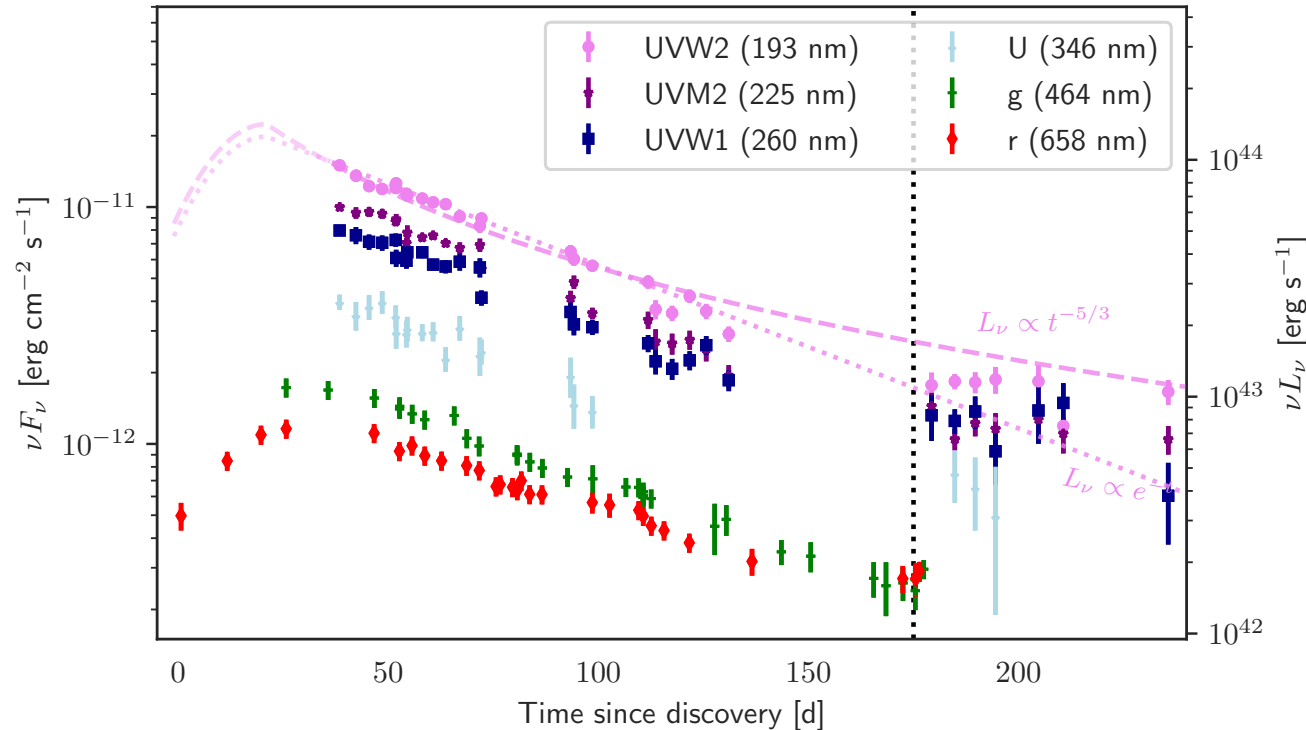
Simeon Reusch *et al.*

Phys. Rev. Lett. **128**, 221101 – Published 3 June 2022

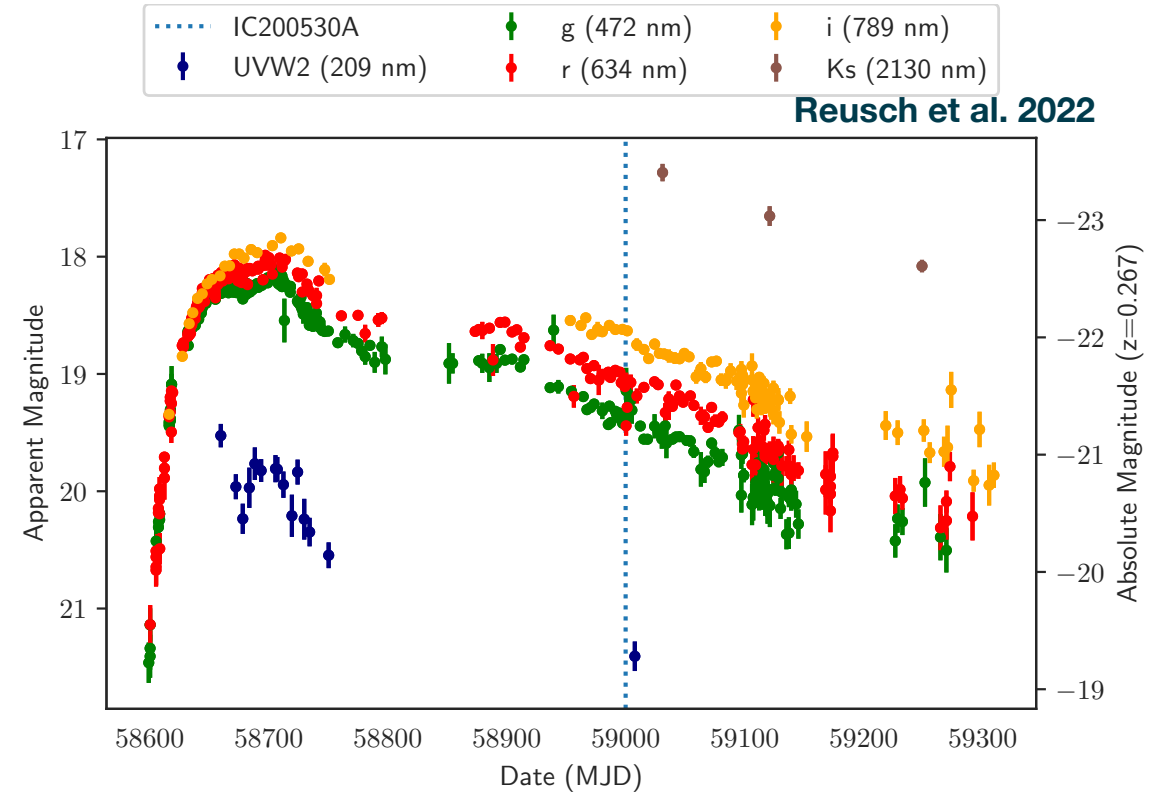
 See Focus story: [Neutrinos from a Black Hole Snack](#)

From 24 neutrino follow-up campaigns -> Now two neutrino-TDEs

Stein et al. 2021



AT2019dsg



AT2019fdr

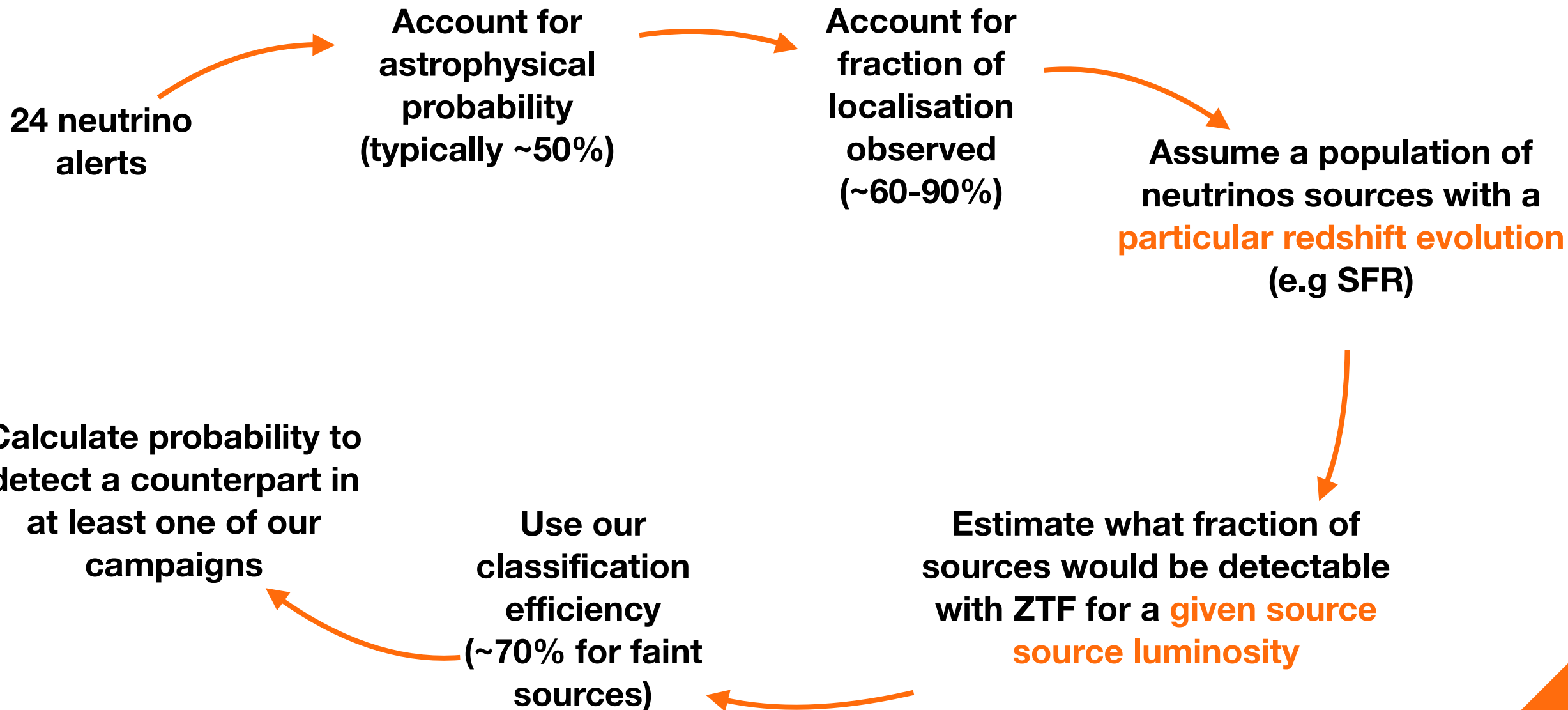
ZTF program uncovered 2 likely neutrino-TDEs out of 24 follow-up campaigns, both particularly bright

Probability of finding two bright TDE by chance with ZTF program: (0.034%=3.4 σ)

See Cecilia's talk from this morning for more info on this

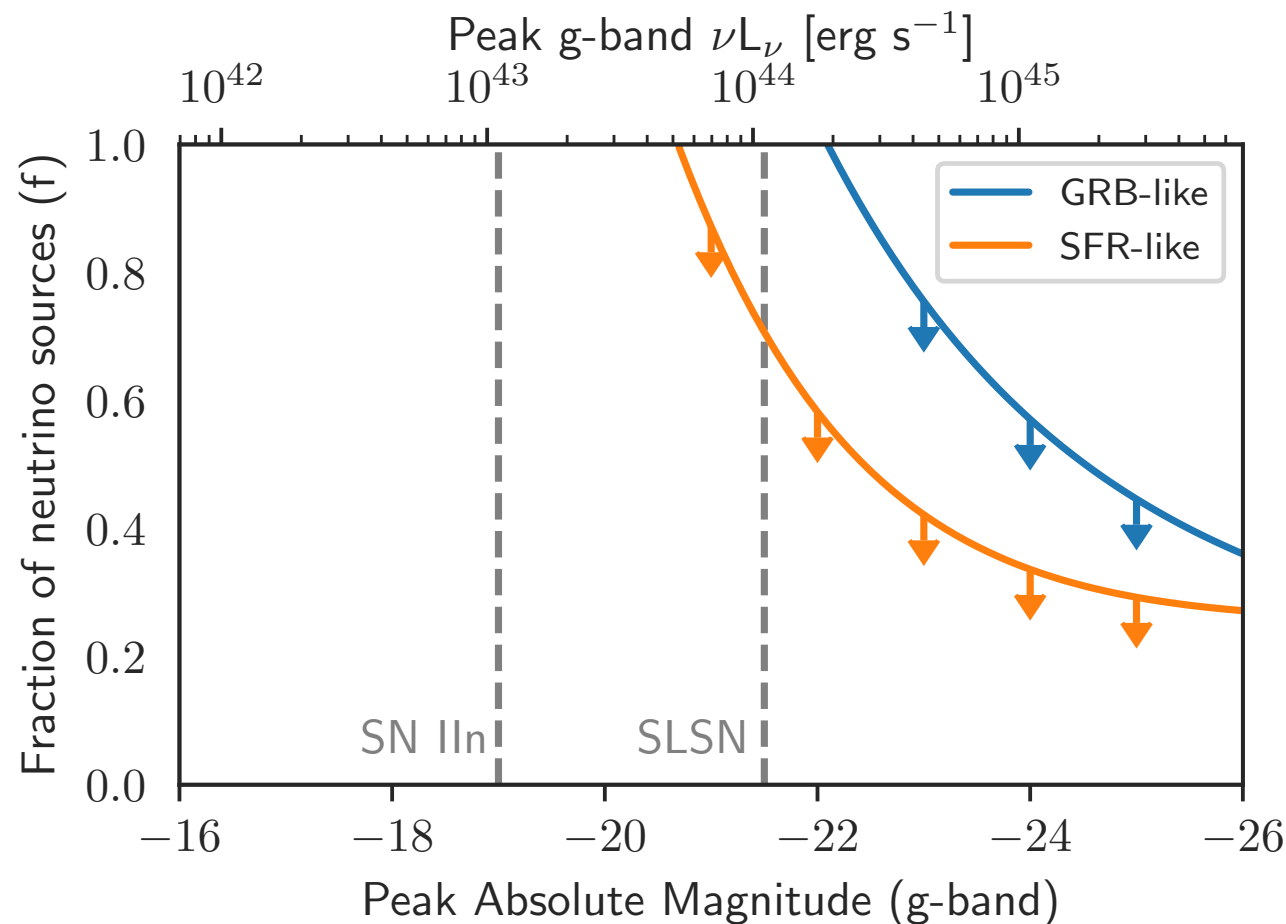
What did we not find?

How likely are you to actually find a counterpart?



Limits on non-TDE neutrino source populations

Stein et al. 2022



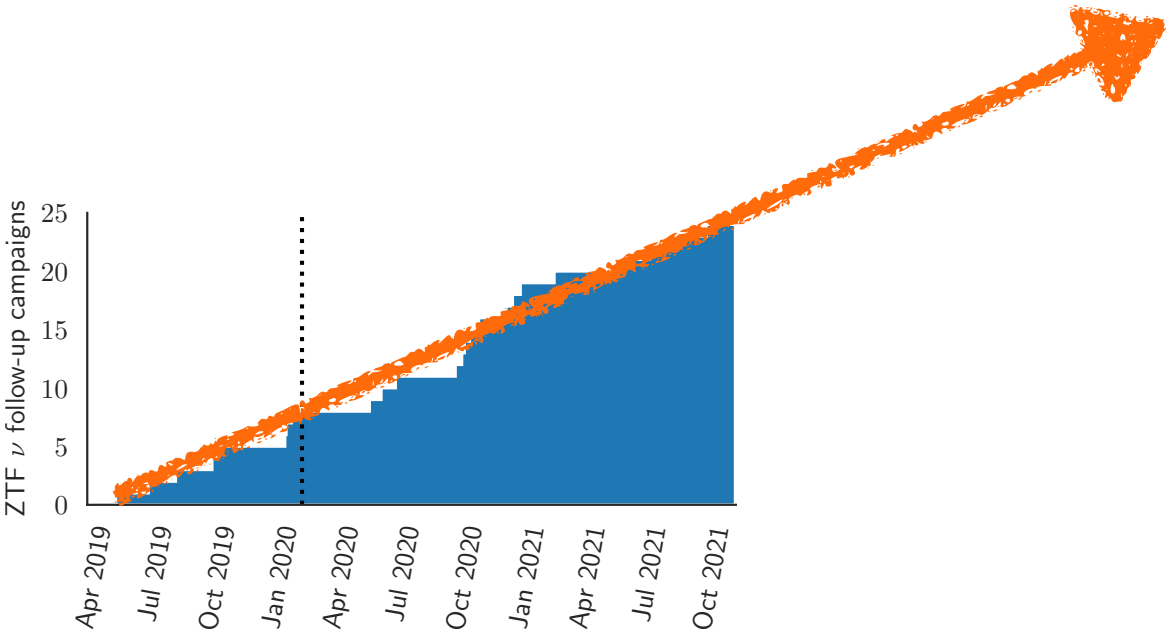
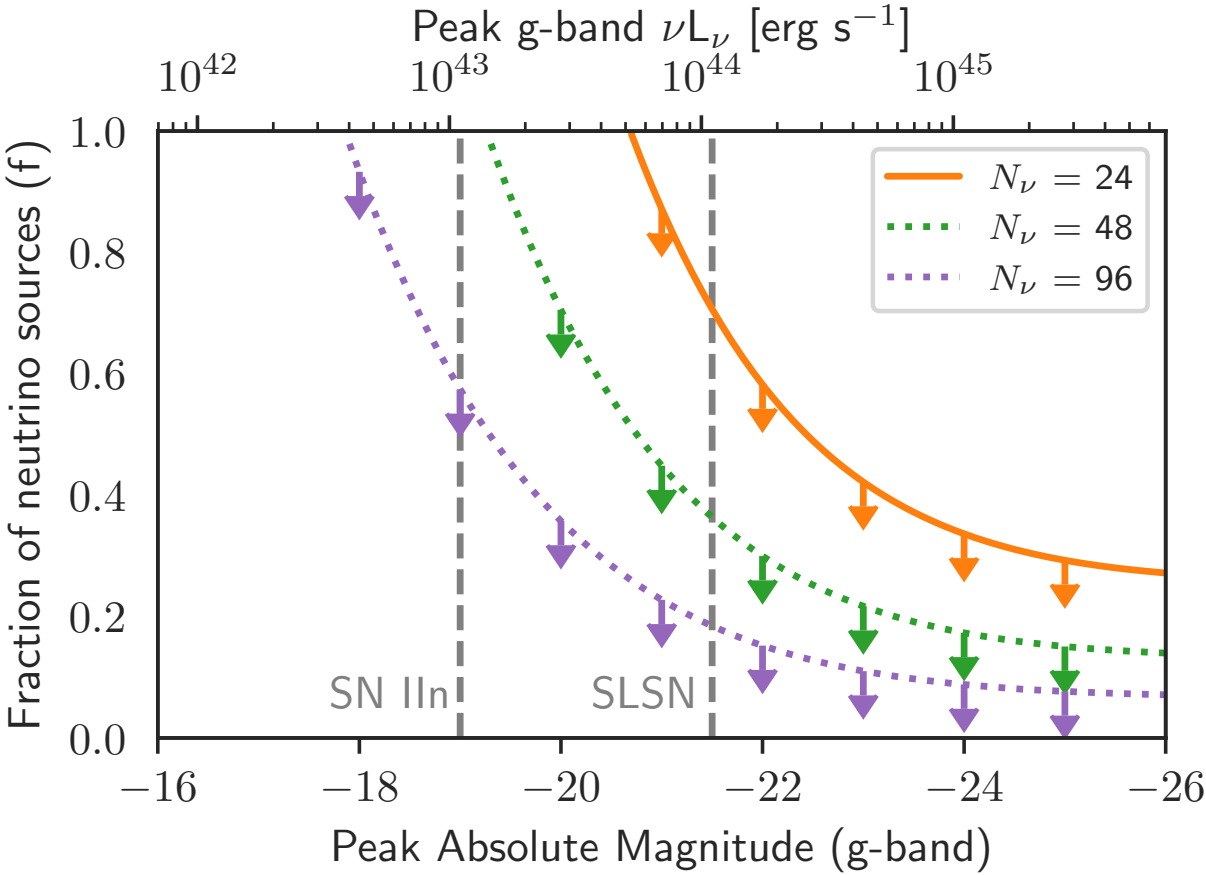
From 24 campaigns -> first constraints on the optical luminosity function of neutrino sources

**How can we do better in
future?**

How can we do better?

Stein et al. 2022

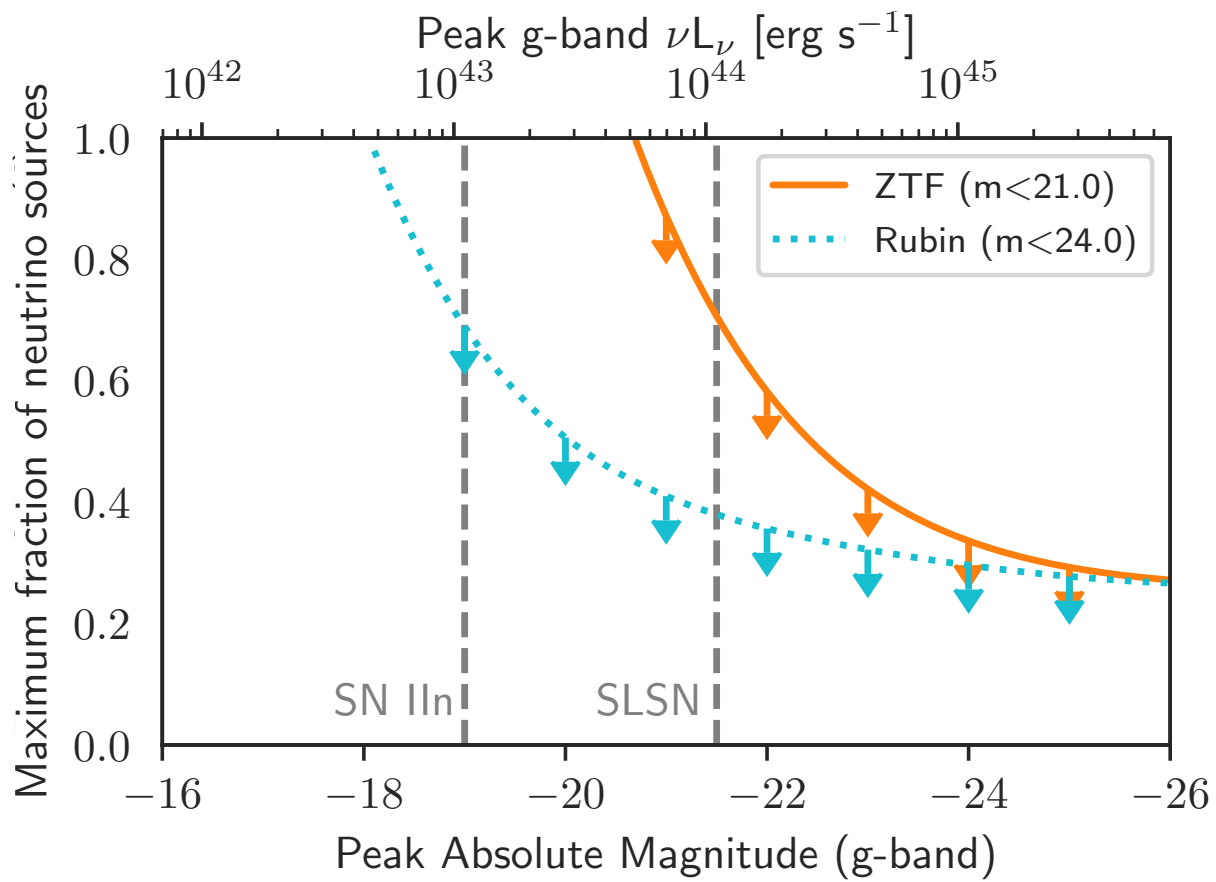
More campaigns



How can we do better?

Stein et al. 2022

Deeper campaigns

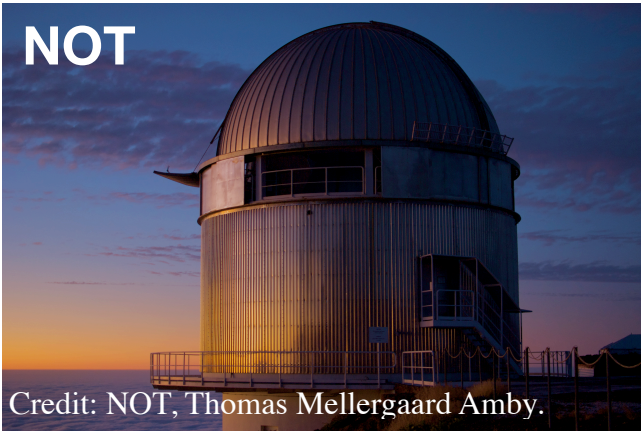
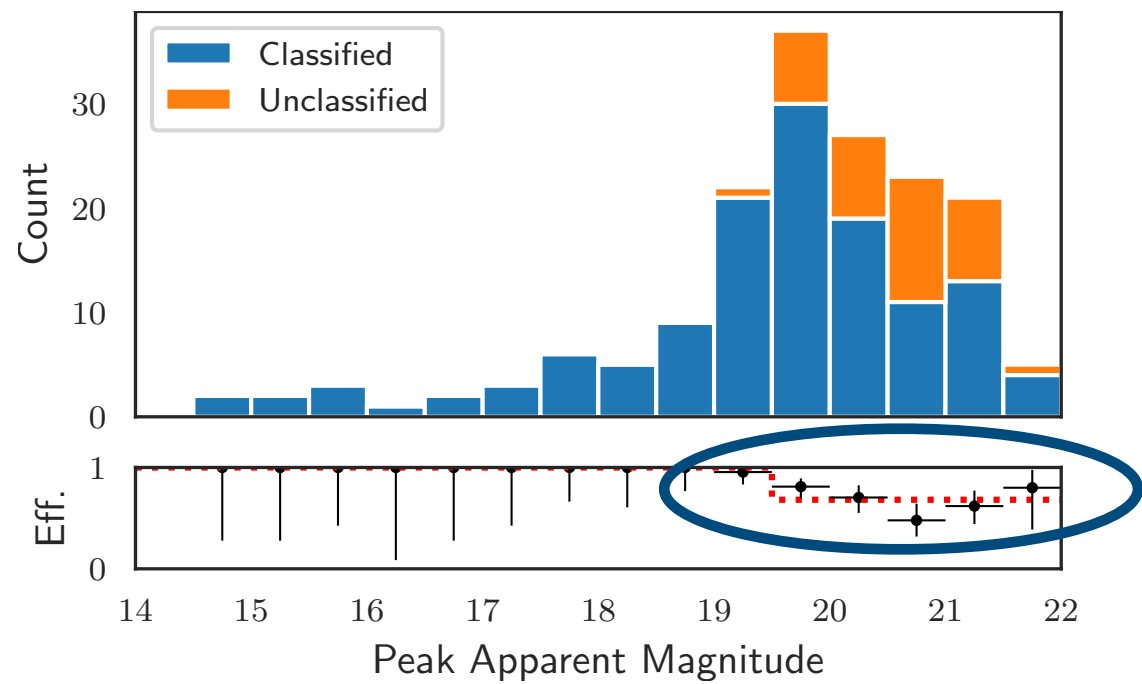


Credit: Bruno C. Quint

How can we do better?

Stein et al. 2022

Better completeness



How can we do better?

New Wavelengths!



Credit: Robert Stein

How can we do better?

New Wavelengths!

13 December 2020

The wide-field infrared transient explorer (WINTER)

Nathan P. Lourie, John W. Baker, Richard S. Burruss, Mark Egan, Gábor Fűrész, Danielle Frostig, Allan A. Garcia-Zych, Nicolae Ganciu, Kari Haworth, Erik Hinrichsen, Mansi M. Kasliwal, Viraj R. Karambelkar, Andrew Malonis, Robert A. Simcoe, Jeffrey Zolkower



Credit: Robert Stein

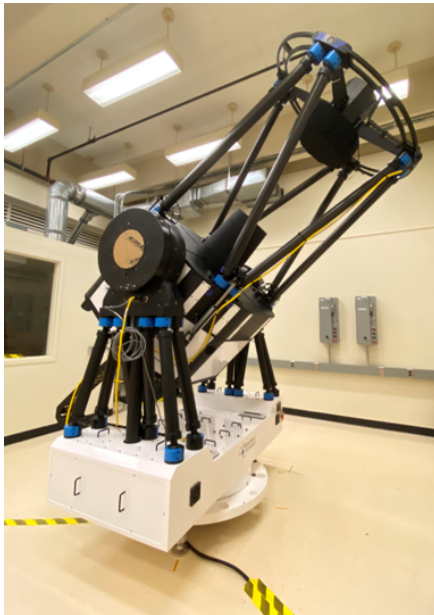
Introducing WINTER

J-band survey with ~monthly cadence

ToO program -> First dedicated IR neutrino follow-up program



Introducing WINTER



J-band survey with ~monthly cadence









ToO program -> First dedicated IR neutrino follow-up program

NIR is also great for kilonova searches ->

<https://arxiv.org/abs/2110.01622>

OPEN ACCESS

An Infrared Search for Kilonovae with the WINTER Telescope. I. Binary Neutron Star Mergers

Danielle Frostig¹ , Sylvia Biscoveanu^{1,2} , Geoffrey Mo^{1,2} , Viraj Karambelkar³ ,
Tito Dal Canton⁴ , Hsin-Yu Chen^{1,2} , Mansi Kasliwal³ , Erik Katsavounidis^{1,2},
Nathan P. Lourie¹, Robert A. Simcoe¹  [+ Show full author list](#)

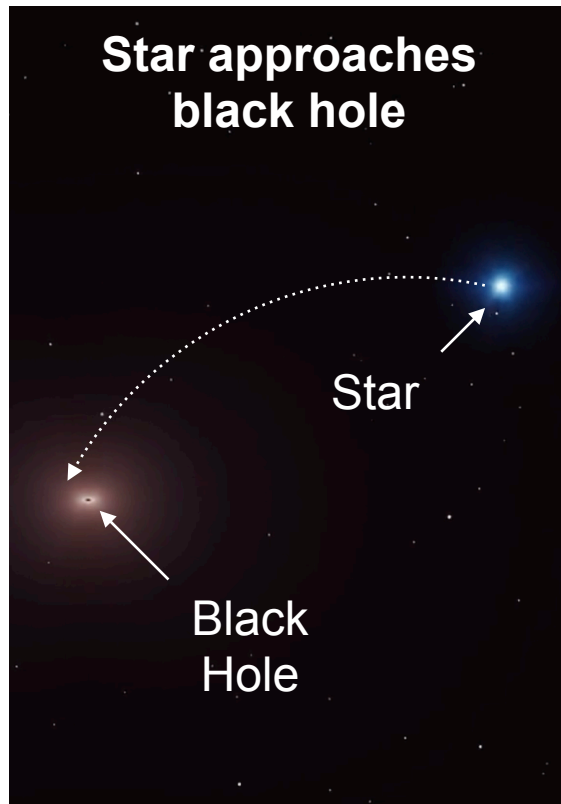
Published 2022 February 21 • © 2022. The Author(s). Published by the American Astronomical Society.

[The Astrophysical Journal](#), [Volume 926](#), [Number 2](#)

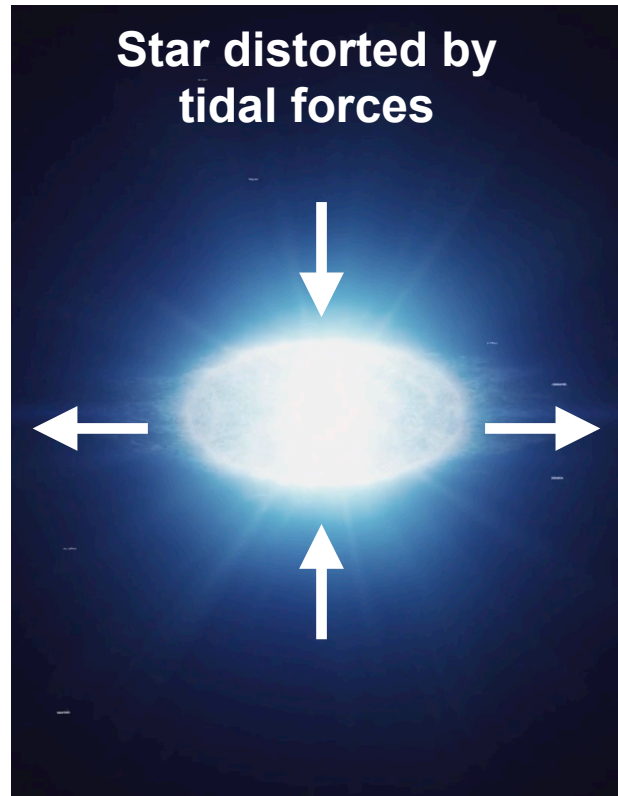
Citation Danielle Frostig *et al* 2022 *ApJ* **926** 152

What are TDEs?

Credit: DESY/Science communication Lab



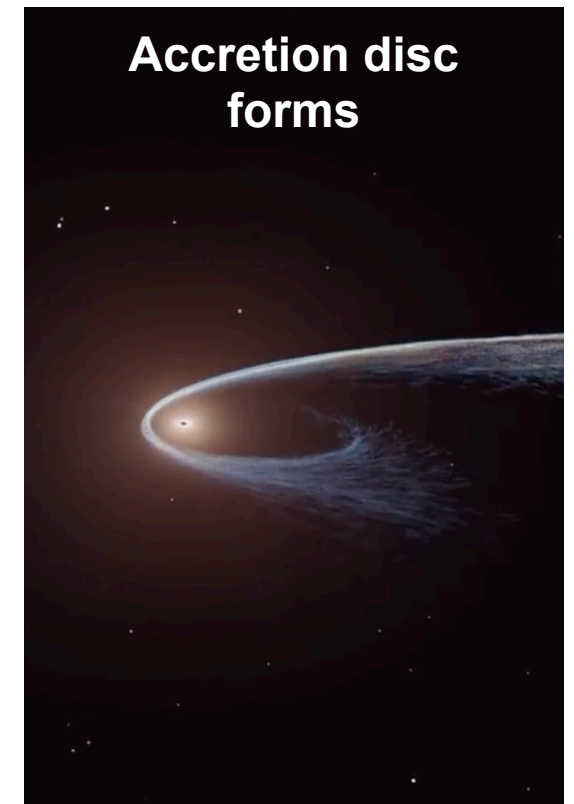
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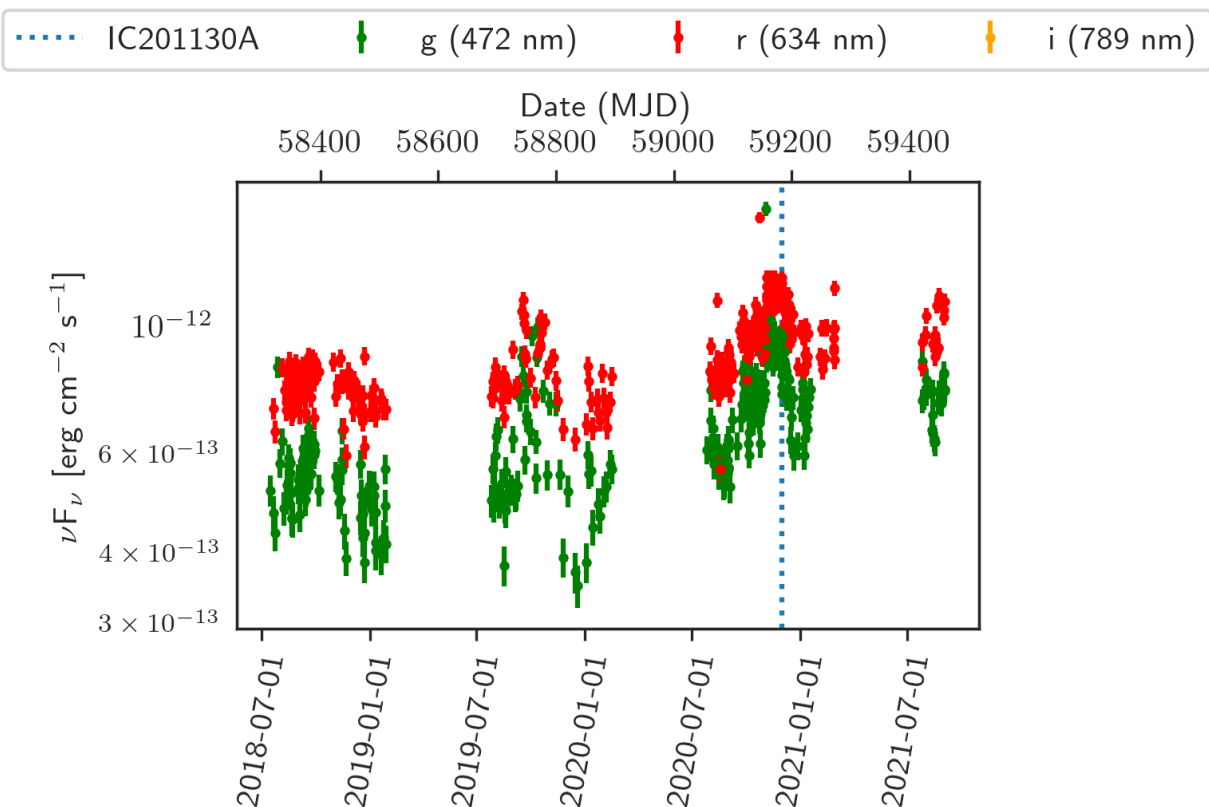


4

Identifying AGN flares

Stein et al. 2022

ZTF Lightcurve of WISEA J020128.20-121946.2



ZTF Lightcurve of WISEA J134034.75+045241.3

